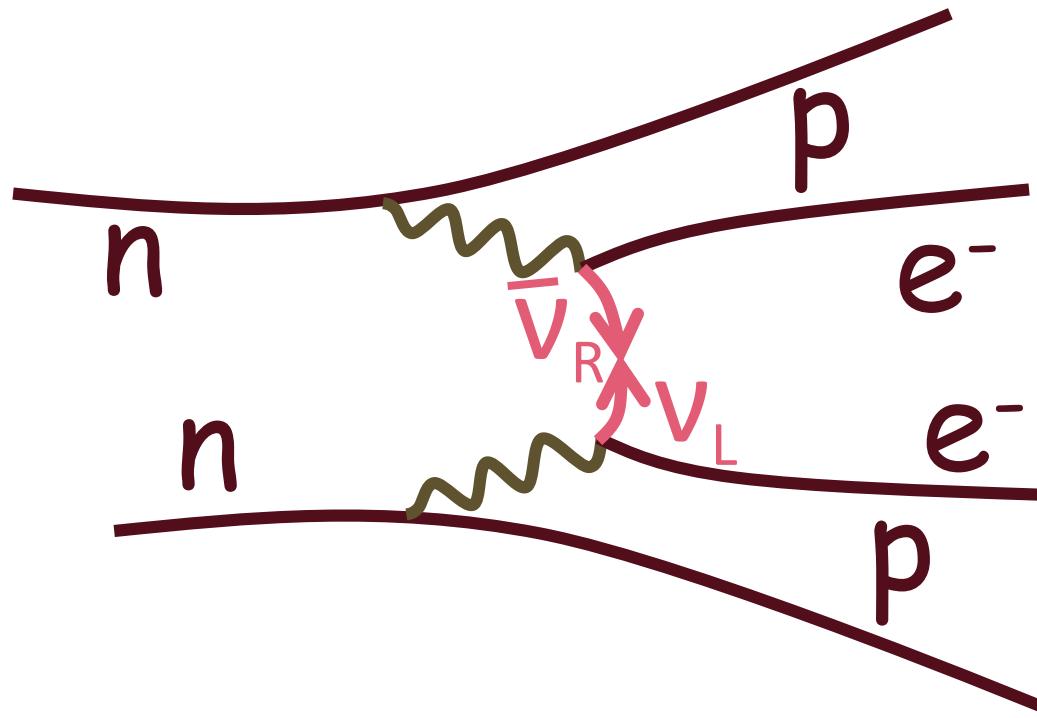
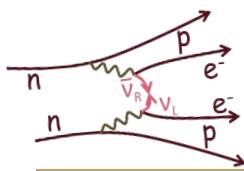


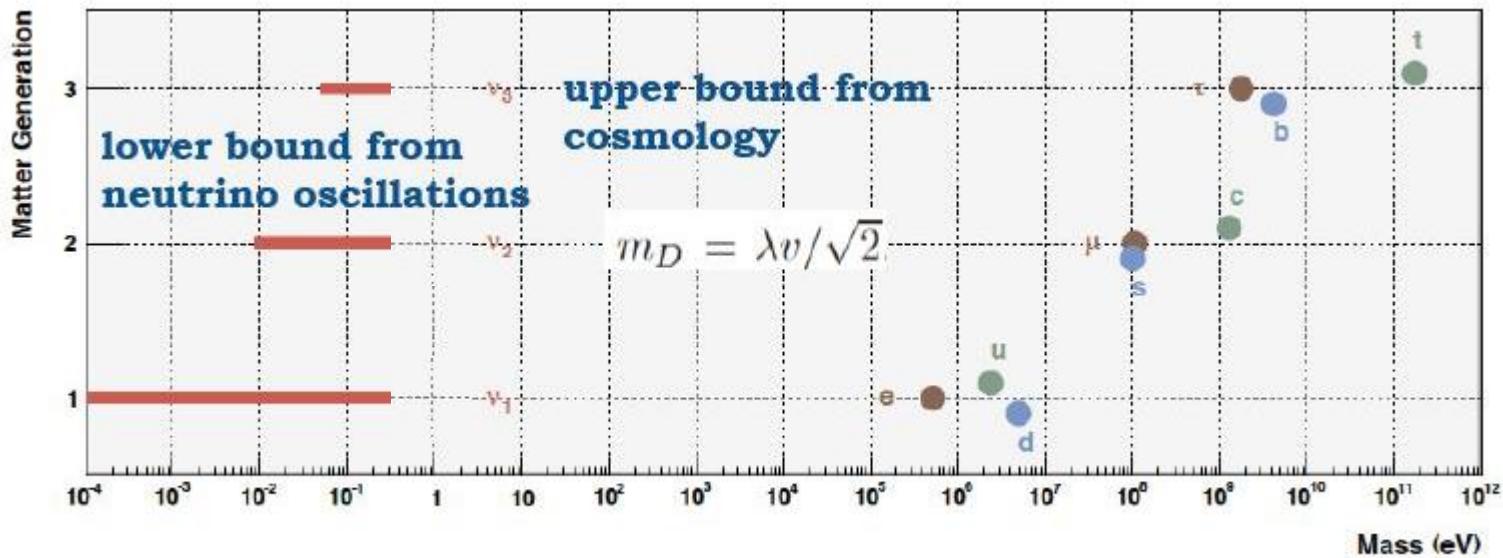
Double Beta Decay and Lepton Number Violation

$$\Delta L \neq 0$$



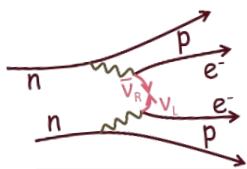


Why are Neutrino Masses so small ?

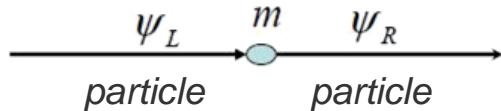


if neutrinos acquire their mass by coupling to the Higgs,
in the same way as the charged leptons,
why should the coupling be so different ?

Neutrino Mass



SM mass terms: $m (\bar{e}_L e_R + \bar{e}_R^c e_L^c)$



$$m_D \bar{\nu}_L \nu_R + m_D \bar{\nu}_L^c \nu_R^c$$

$$= \bar{N}_L \begin{pmatrix} 0 & m_D \\ m_D & 0 \end{pmatrix} N_R$$

two degenerate mass-eigenvalues m_D

no ν_R or ν_L^c in standard model

\Rightarrow no \mathcal{V} -massterm

\Rightarrow add ν_R , ν_L^c

$$N_L = \begin{pmatrix} \nu_L \\ \nu_L^c \end{pmatrix}$$

$$N_R = \begin{pmatrix} \nu_R^c \\ \nu_R \end{pmatrix}$$

(active
sterile)

Dirac

$$\nu_L \quad \nu_L^c$$



$$\nu_R^c \quad \nu_R$$

$$\mathbf{v}_D \neq \mathbf{v}_D^c$$

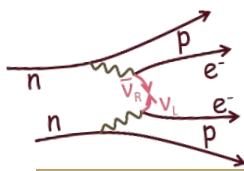
Majorana

$$\nu_{M,L} = \nu_L + \nu_L^c$$



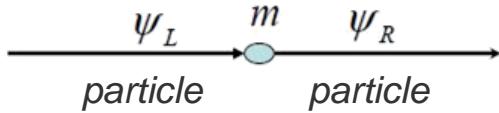
$$\nu_{M,R} = \nu_R^c + \nu_R$$

$$\mathbf{v}_M = \mathbf{v}_M^c$$



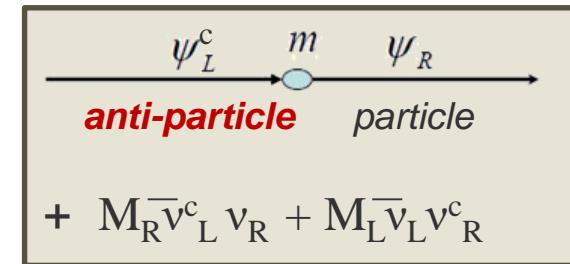
Majorana Neutrino

$$\mathbf{v}_M = \mathbf{v}_M^c$$



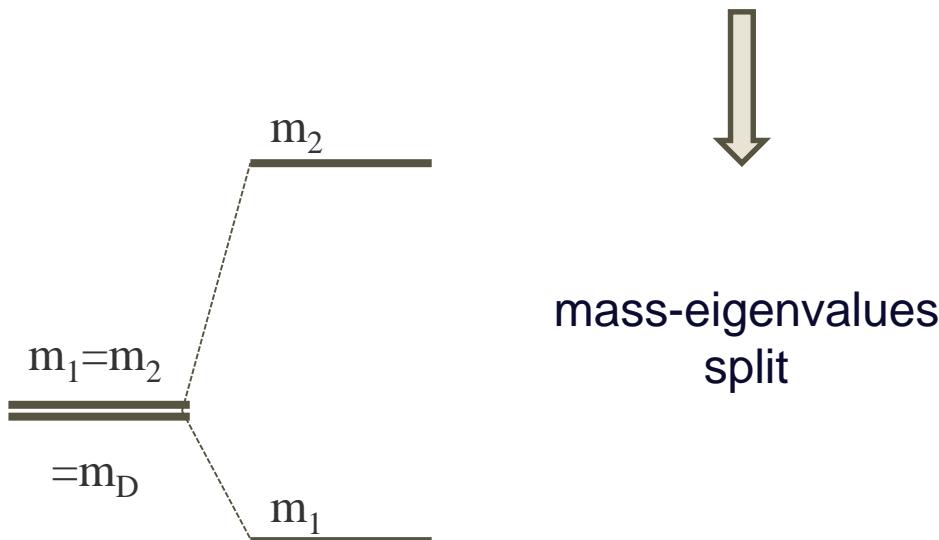
$$m_D \bar{v}_L v_R + m_D \bar{v}_L^c v_R^c$$

$$= \overline{N}_L \begin{pmatrix} M_L & m_D \\ m_D & M_R \end{pmatrix} N_R$$

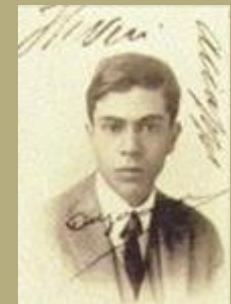


$$N_L = \begin{pmatrix} v_L \\ v_L^c \end{pmatrix}$$

$$N_R = \begin{pmatrix} v_R^c \\ v_R \end{pmatrix}$$

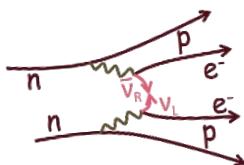


Majorana



$v_{M,L} = v_L + v_L^c$
 $v_{M,R} = v_R^c + v_R$

$\mathbf{v}_M = \mathbf{v}_M^c$



Majorana Neutrino

$$\mathbf{v}_M = \mathbf{v}_M^c$$

$$\begin{pmatrix} 0 & m_D \\ m_D & M_R \end{pmatrix}$$

m_D normal Fermion-mass
 $M_R \sim$ higher scale (GUT ?)

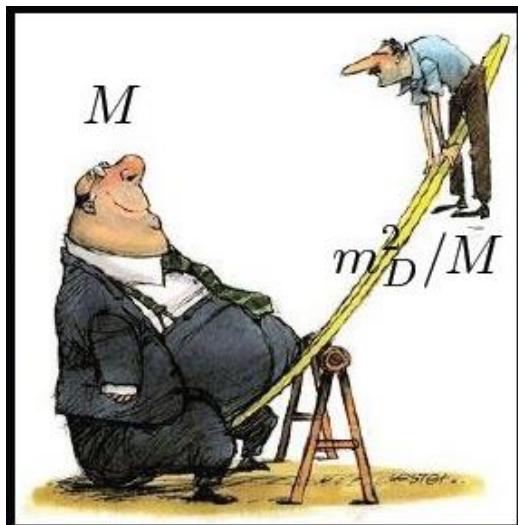
$$m_1 = m_D^2 / M_R$$

mostly left-handed, active

=> \mathbf{v} we know

$$m_2 = M_R$$

mostly right-handed, sterile



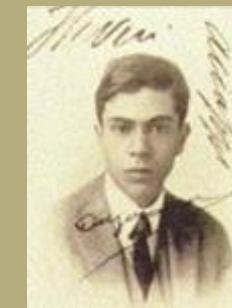
to get
 $m_1 \sim \text{meV}$

$$\downarrow$$

$M_R \sim 10^{15} \text{ eV}$

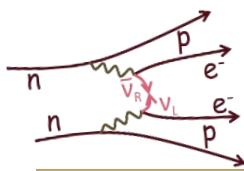
Majorana

$$v_{M,L} = v_L + v_L^c$$



$$v_{M,R} = v_R^c + v_R$$

$$\mathbf{v}_M = \mathbf{v}_M^c$$



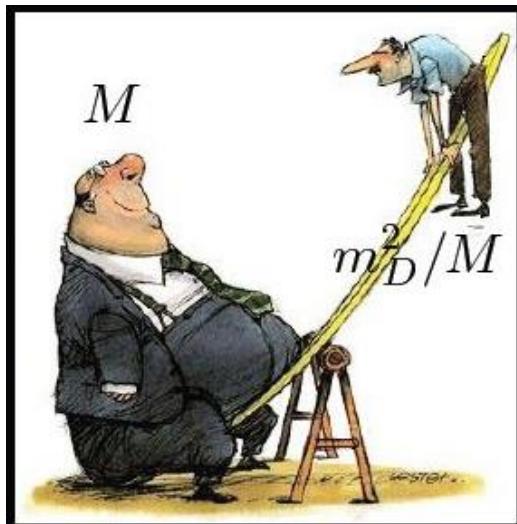
Majorana Neutrino

$$\mathbf{v}_M = \mathbf{v}_M^c$$

- \mathbf{v} has no charge \Rightarrow can be Majorana-particle
 \Rightarrow can explain small \mathbf{v} – mass (even with normal m_D)

Neutrino is a normal Fermion, it just happen to have no charge

- \Rightarrow heavy right handed partner
- \Rightarrow decays in early Universe $\Delta L \neq 0$
- \Rightarrow creates Matter - Antimatter imbalance



to get
 $m_1 \sim \text{meV}$

↓

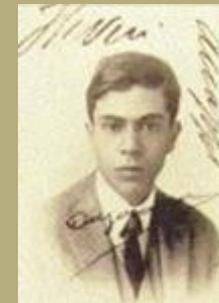
$M_R \sim 10^{15} \text{ eV}$

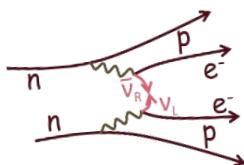
Majorana

$v_{M,L} = v_L + v_L^c$

$v_{M,R} = v_R^c + v_R$

$\mathbf{v}_M = \mathbf{v}_M^c$





Majorana Neutrino

$$\mathbf{v}_M = \mathbf{v}_M^c$$

\mathbf{v} has no charge \Rightarrow can be Majorana-particle

\Rightarrow can explain small \mathbf{v} – mass (even with normal m_D)

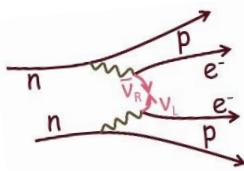
Neutrino is a normal Fermion, it just happen to have no charge

\Rightarrow heavy right handed partner

\Rightarrow decays in early Universe $\Delta L \neq 0$

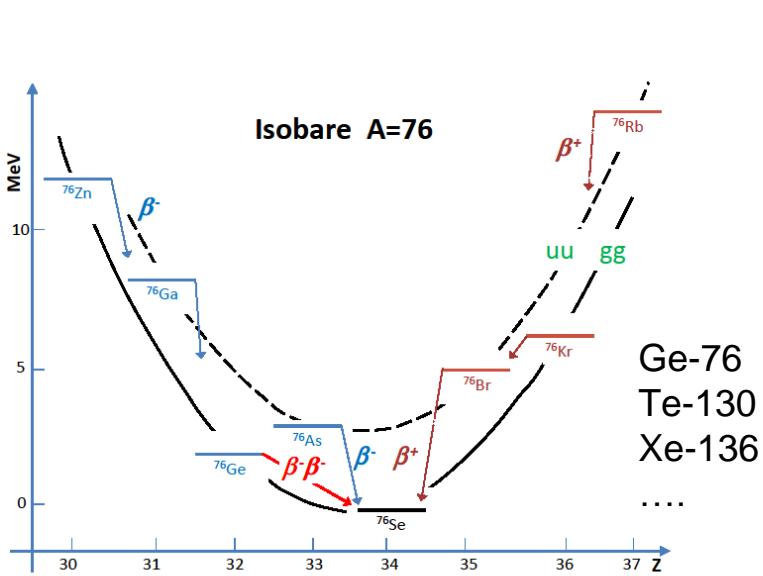
\Rightarrow creates Matter - Antimatter imbalance





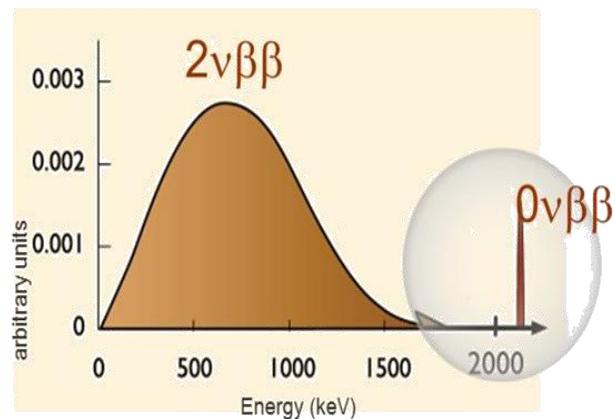
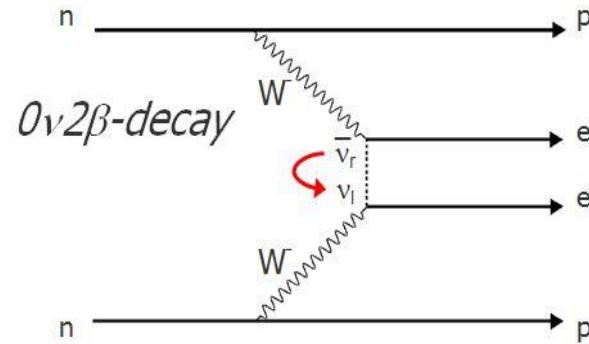
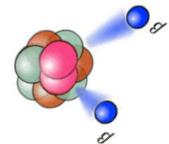
Search for Neutrino-less Double Beta Decay

$$\Delta L \neq 0$$

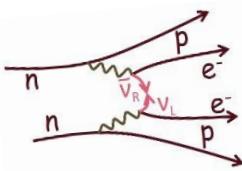


easiest but not easy way to see

if ν are Majorana-type



⇒ checks
Majorana character
⇒ very sensitive to m_ν



Sensitivity

$\Delta L \neq 0$

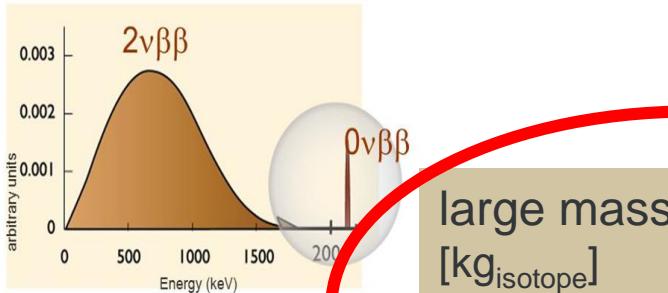
effective neutrino mass

$$1 / T_{1/2}^{0\nu} = G \cdot NME^2 \cdot m_{\beta\beta}^2$$

phase space nuclear matrix element
 $\sim Q^5$

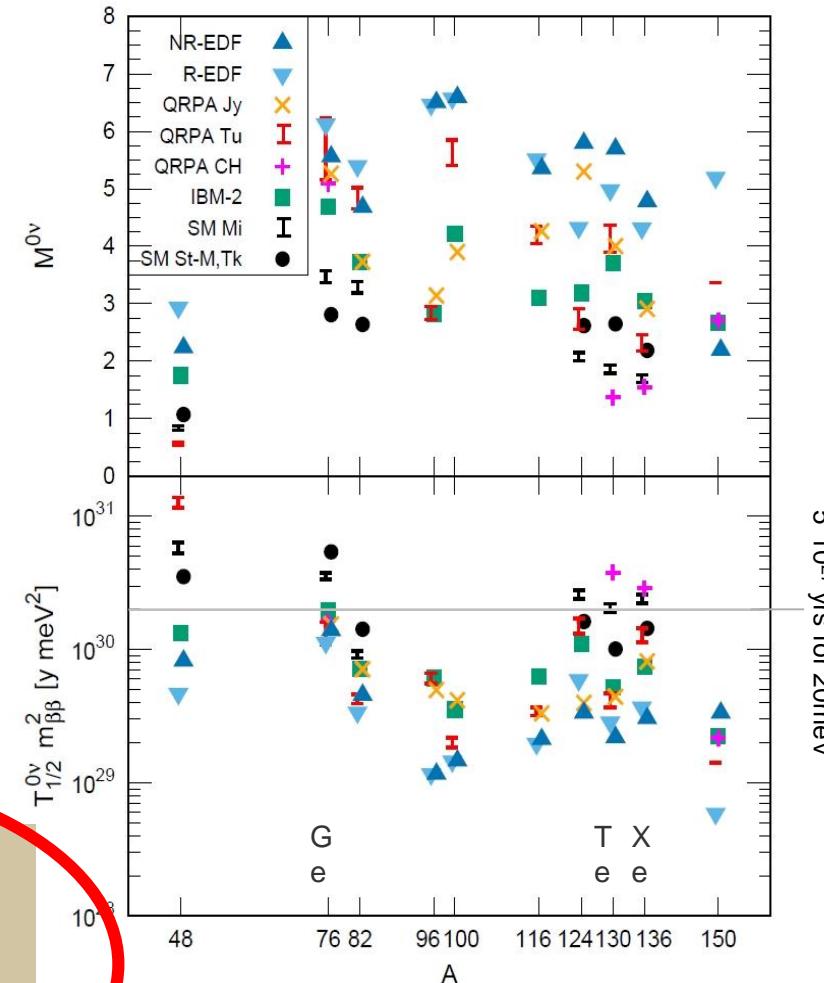
no favored isotope
 considering spread of nuclear
 matrix elements and Q-values

NME extremely important to get $m_{\beta\beta}^2$

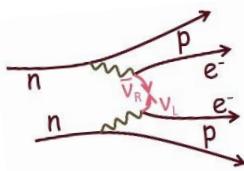


large mass
 $[kg_{isotope}]$

low background in ROI
 $[cts / FWHM t_{isotope} yr]$



arXiv:
 1610.06548v1



Sensitivity

$$\Delta L \neq 0$$

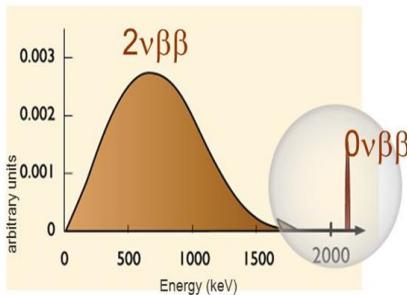
effective neutrino mass

$$1 / T_{1/2}^{0\nu} = G \cdot N M E^2 \cdot m_{\beta\beta}^2$$

phase space nuclear matrix element
 $\sim Q^5$

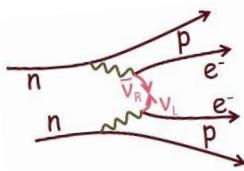
sensitivity on $T_{1/2}^{0\nu}$

mid term: a few 10^{26} yrs ($m_{\beta\beta} \sim 40-100$ meV)
 long term: a few 10^{27} yrs ($m_{\beta\beta} \sim 10-20$ meV)

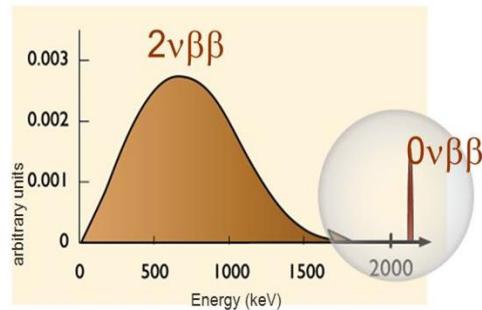


large mass
 $[kg_{isotope}]$

low background in ROI
 $[cts / FWHM t_{isotope} yr]$



Double Beta Decay and Lepton Number Violation $\Delta L \neq 0$



sensitivity on $T^{0\nu}_{1/2}$

easiest but not easy way to see
if ν are Majorana-type

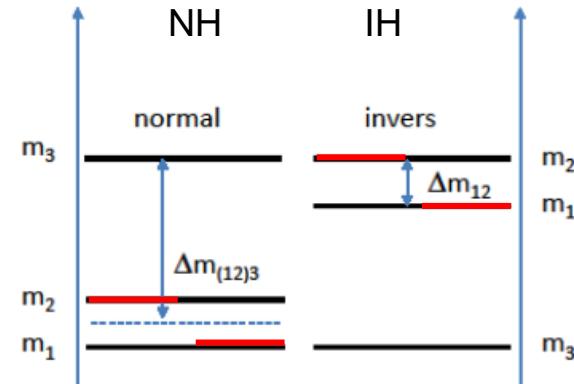
mid term: a few 10^{26} yrs ($m_{\beta\beta} \sim 40-100$ meV)
long term: a few 10^{27} yrs ($m_{\beta\beta} \sim 10-20$ meV)

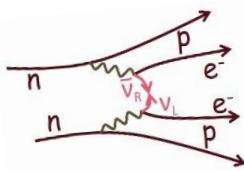
via ν exchange

effective neutrino mass

$$1 / T^{0\nu}_{1/2} = G \cdot NME^2 \cdot m_{\beta\beta}^2$$

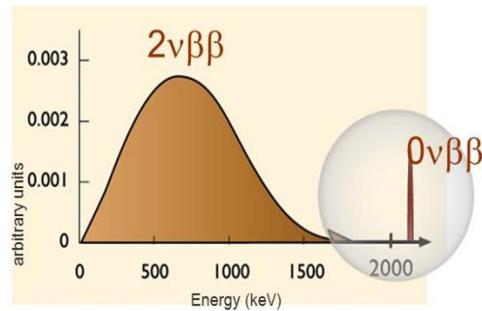
phase space nuclear matrix element





Double Beta Decay and Lepton Number Violation

$$\Delta L \neq 0$$

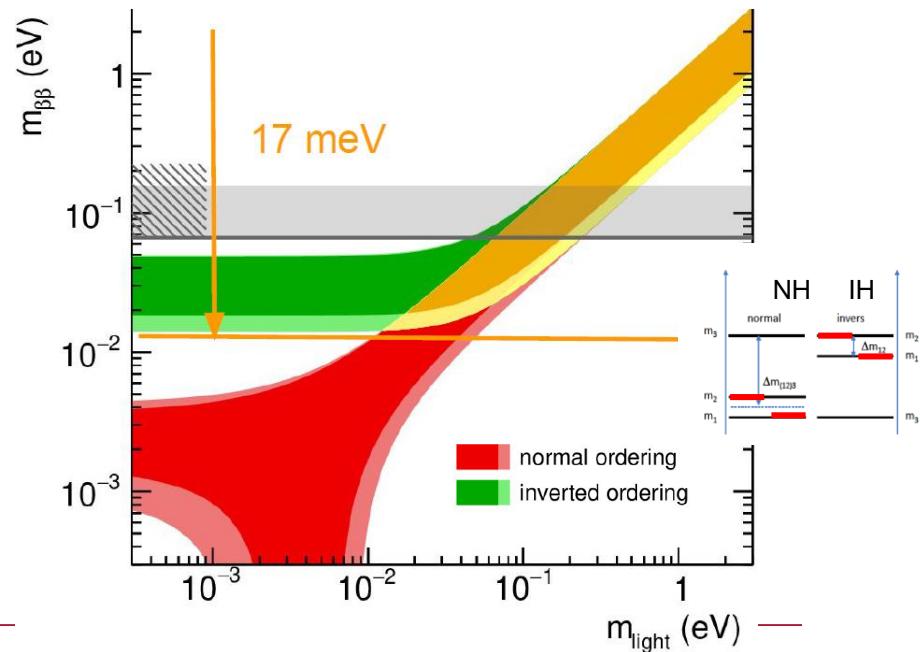
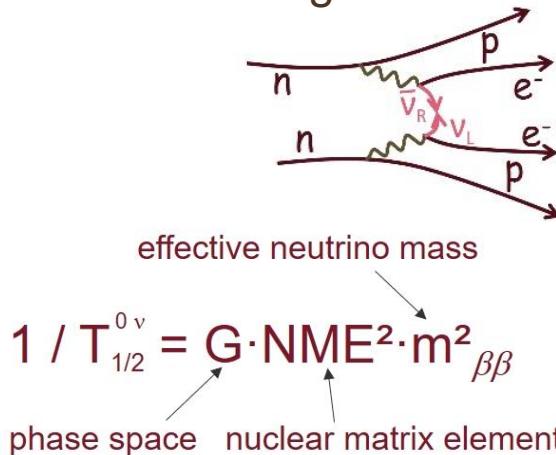


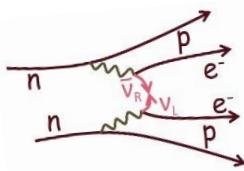
sensitivity on $T^{0\nu}_{1/2}$

easiest but not easy way to see
if ν are Majorana-type

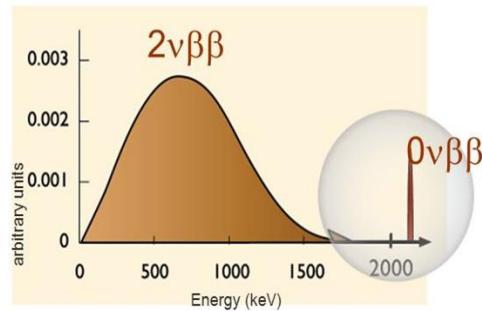
mid term: a few 10^{26} yrs ($m_{\beta\beta} \sim 40-100$ meV)
long term: a few 10^{27} yrs ($m_{\beta\beta} \sim 10-20$ meV)

via ν exchange





Double Beta Decay and Lepton Number Violation $\Delta L \neq 0$

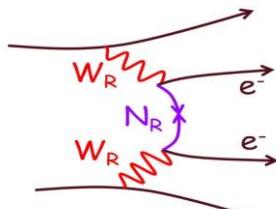


sensitivity on $T^{0\nu}_{1/2}$

easiest but not easy way to see
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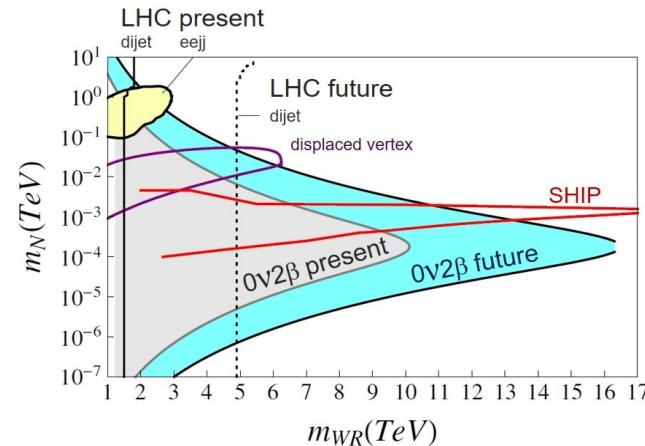
mid term: a few 10^{26} yrs ($m_{\beta\beta} \sim 40-100$ meV)
long term: a few 10^{27} yrs ($m_{\beta\beta} \sim 10-20$ meV)

other $\Delta L \neq 0$ processes

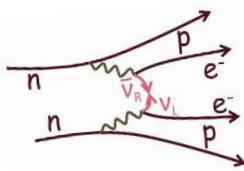


LR symmetry

heavy W_R and N_R exchange



$0\nu\beta\beta$ searches complementary and competitive
- $\Delta L \neq 0$ is the key, not so much $m_{\beta\beta}$ -



Search for Neutrino-less Double Beta Decay

$\Delta L \neq 0$

Ge detectors

GERDA

very good ΔE (narrow ROI)

Majorana

LEGEND

liquid noble gas Xe

EXO

large detector masses

nEXO

loaded liquid scintillator Xe, Te

KamLAND-Zen

self shielding

KamLAND2-Zen

SNO+

gaseous detectors Xe, Se,Nd,Ca

SuperNEMO

NEXT

tracking

PANDA - X

CdZnTe detectors

COBRA

cryo bolometers Te

CUORE

larger variety of isotopes
new techniques

cryo + light Te, Mo

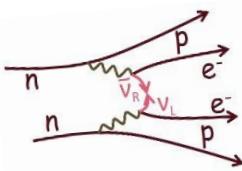
Cupid, AMoRe

competitive limits

running

in preparation

R&D and future projects



Liquid Noble Gases - EXO 200 / nEXO

$\Delta L \neq 0$

liquid Xenon TPC enriched in ^{136}Xe (80.6 %), charge and light detection

WIPP New Mexico USA

$\Delta E \sim 70 \text{ keV FWHM}$

self shielding, multi site recognition

EXO 200: 170 kg_{isotope} total
/ 80 kg_{isotope} active volume

results:

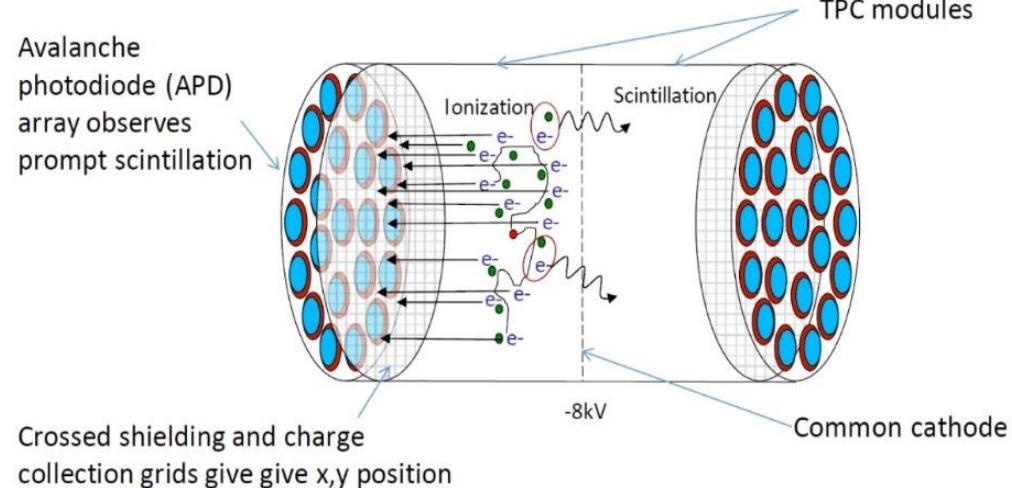
234 kg·yr exposure

sensitivity $5.0 \cdot 10^{25} \text{ yr}$
 $T_{1/2}^{0\nu\beta\beta} > 3.5 \cdot 10^{25} \text{ yr}$

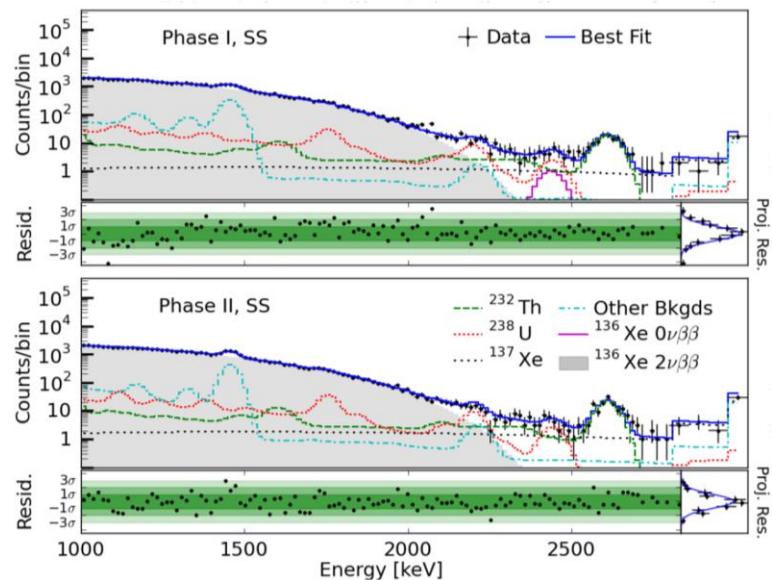
background in ROI
 $170 / \text{FWHM} \cdot t_{\text{isotope}} \cdot \text{yr}$

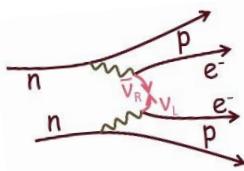
PRL 120 072701 (2018)

arXiv 1906.02723



Crossed shielding and charge collection grids give x,y position





Liquid Noble Gases - nEXO

$$\Delta L \neq 0$$

liquid Xenon single TPC

5000 kg enriched IXe

expected to be at **SNOLAB**

improved performance:

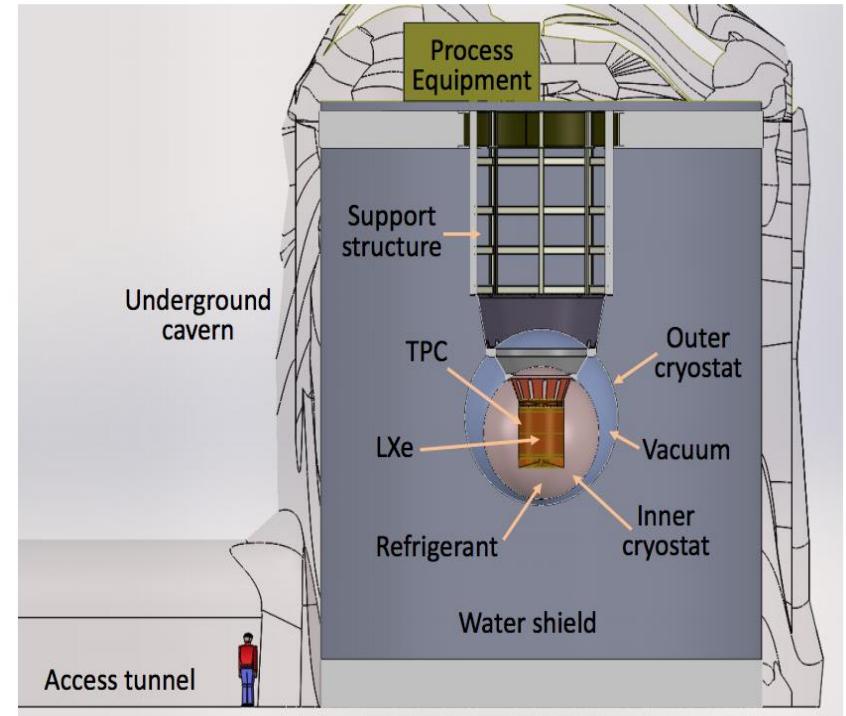
energy and position resolution

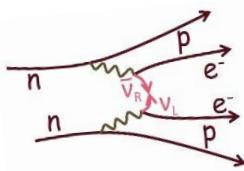
R&D on Ba-tagging

goal $T_{1/2}^{0\nu\beta\beta} > 9.2 \cdot 10^{27}$ yr exclusion
 $5.7 \cdot 10^{27}$ yr discovery

background in ROI

$$\sim 0.6 / \text{FWHM} \cdot t_{\text{isotope}} \cdot \text{yr}$$





Liquid Scintillator - KamLAND-Zen

$\Delta L \neq 0$

3 m diam. balloon: **liquid scintillator** loaded with enriched Xenon inserted into KamLAND

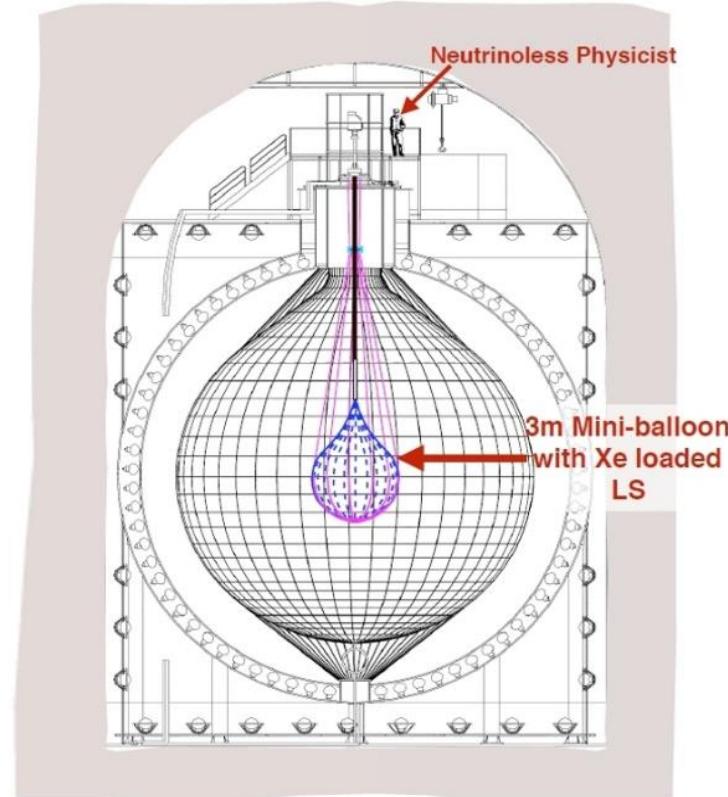
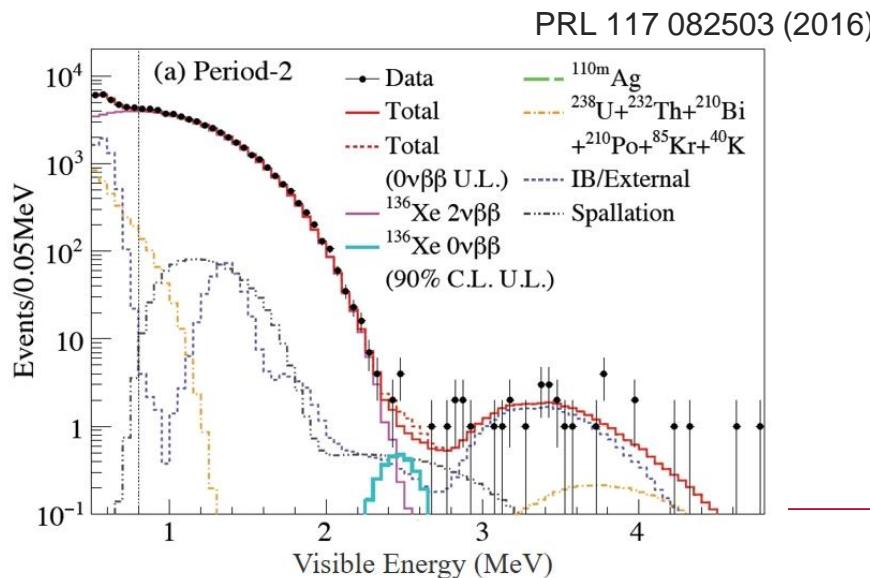
$\Delta E \sim 250$ keV FWHM

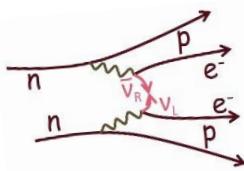
results: 383 kg Xe / 110 kg_{isotope} in FV
 ~ 600 kg·yr

$$T_{1/2}^{0\nu\beta\beta} > 10.7 \cdot 10^{25} \text{ yr}$$

$$\text{sensitivity } 5.6 \cdot 10^{25} \text{ yr}$$

$$\text{background in ROI } \sim 60 / \text{FWHM} \cdot t_{\text{isotope}} \cdot \text{yr}$$





Liquid Scintillator - SNO+

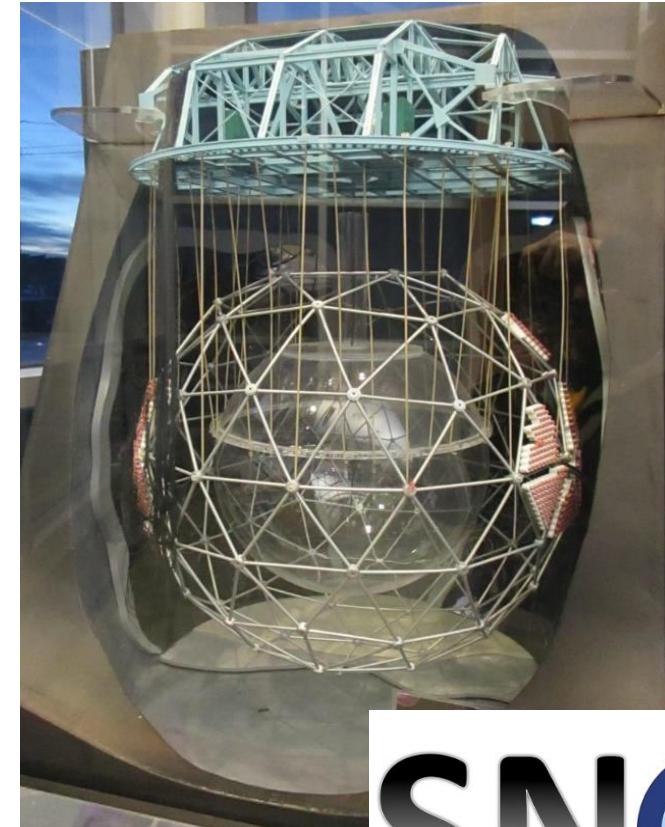
$\Delta L \neq 0$

infrastructure at SNOLAB

Phase I:
acrylic vessel filled with LS
+ 4000 kg of nat-Te
 $(\sim 30\% \text{ Te-130})$

LS filling 2019
 Te loading 2020

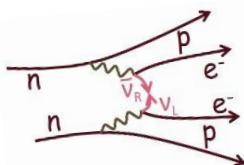
sensitivity goal $T_{1/2}^{0\nu\beta\beta} > 1.9 \cdot 10^{26} \text{ yr}$



Phase II: more Te, better FWHM

THEIA project: 50 kton water-based liquid scintillator
 solar-v will be dominant background





Calorimetry - Cuore @ LNGS

$\Delta L \neq 0$

calorimetry at mK temperature
in natural TeO_2 crystals ^{130}Te (30%)

CUORE 750 kg TeO_2 (206 kg ^{130}Te)
988 crystals

results 2019:

369,9 kg.yr (103 kg·yr ^{130}Te)

$\Delta E \sim 8 \text{ keV FWHM}$

$$T_{1/2}^{0\nu\beta\beta} > 2,3 \cdot 10^{25} \text{ yr}$$

background in ROI

$$\sim 450 / \text{FWHM} \cdot t_{\text{isotope}} \cdot \text{yr}$$

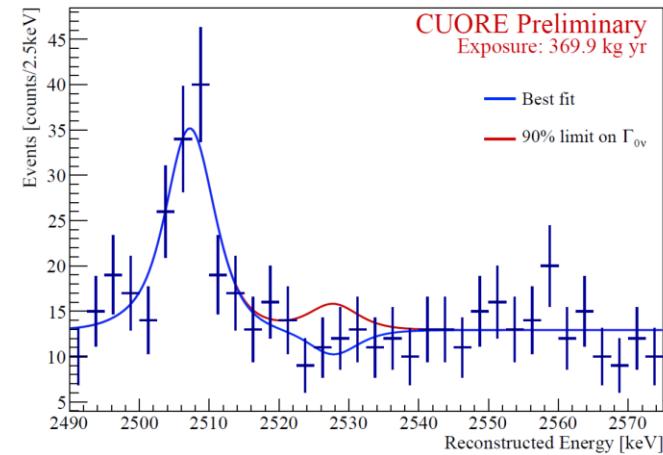
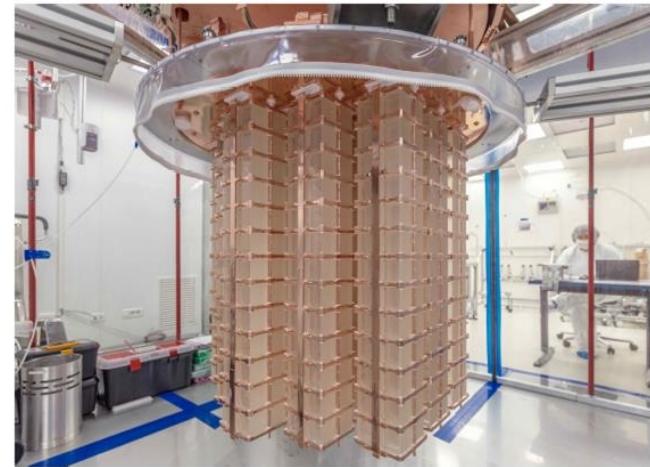
goals:

goal sensitivity

$$T_{1/2}^{0\nu\beta\beta} > 9.5 \cdot 10^{25} \text{ yr}$$

background

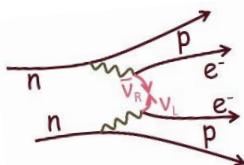
$$\sim 180 / \text{FWHM} \cdot t_{\text{isotope}} \cdot \text{yr}$$



Phys.Rev.Lett 120, 132501 (2018)
Update: TAUP 2019 conference

potential upgrade with calorimetry + light

CUORE => CUPID



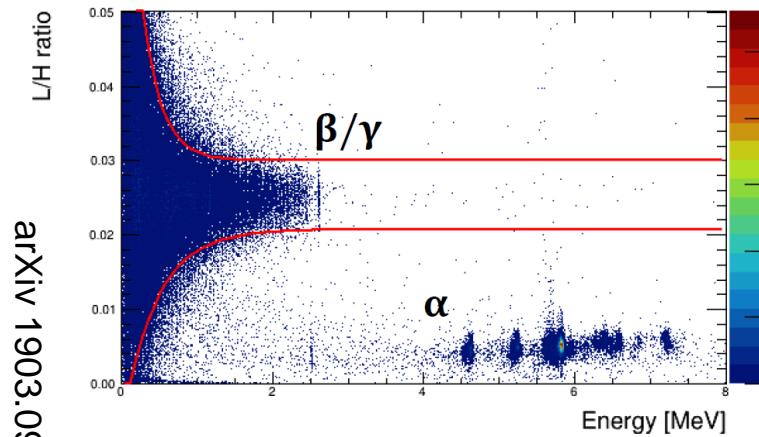
Scintillating bolometers

$$\Delta L \neq 0$$

AMORE @ Yangyang UGL Korea
goal

sensitivity $T_{1/2}^{0\nu\beta\beta} > 3 \cdot 10^{26}$ yr
for ^{100}Mo 250kg·yr exposure

AMORE-Pilot results
2kg CaMoO₄, 0,3 kg*yrs exposure
still very high background
 $T_{1/2}^{0\nu\beta\beta} > 9,5 \cdot 10^{22}$ yr



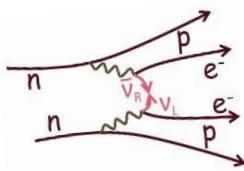
arXiv 1903.09483

CANDLES @ Kamioka / Japan
 ^{48}Ca hightes Q-value 4,27MeV, lowest abundance 0,187%

305 kg CaF₂ scintillator detectors
in liquid scintillator
result from 131 days
 $T_{1/2}^{0\nu\beta\beta} > 6,2 \cdot 10^{22}$ yr

working on scintillating bolometers
first demonstration results





Ge Semiconductor - MAJORANA

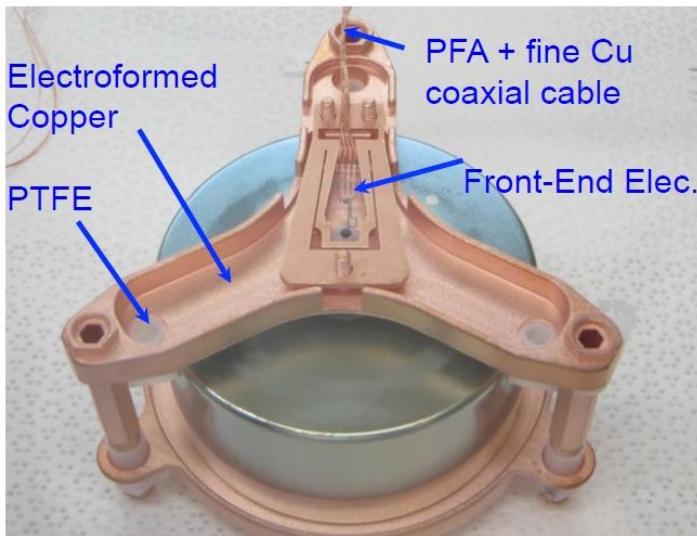
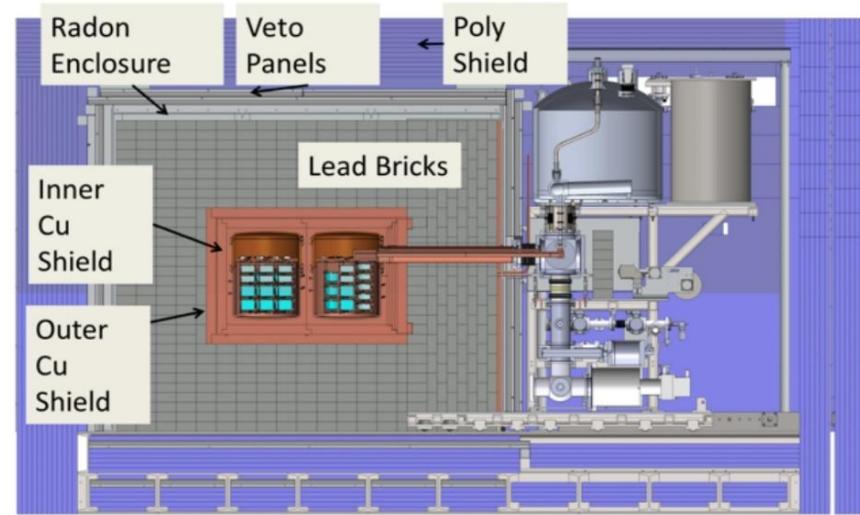
$$\Delta L \neq 0$$

Ge detectors

Cu/Pb shielding

Sanford Lab , USA

44.1 kg ^{76}Ge (88%) running



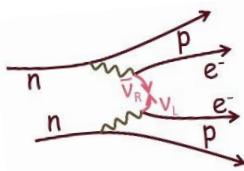
PRC 100 025501 (2019)

results

26 kg·yr

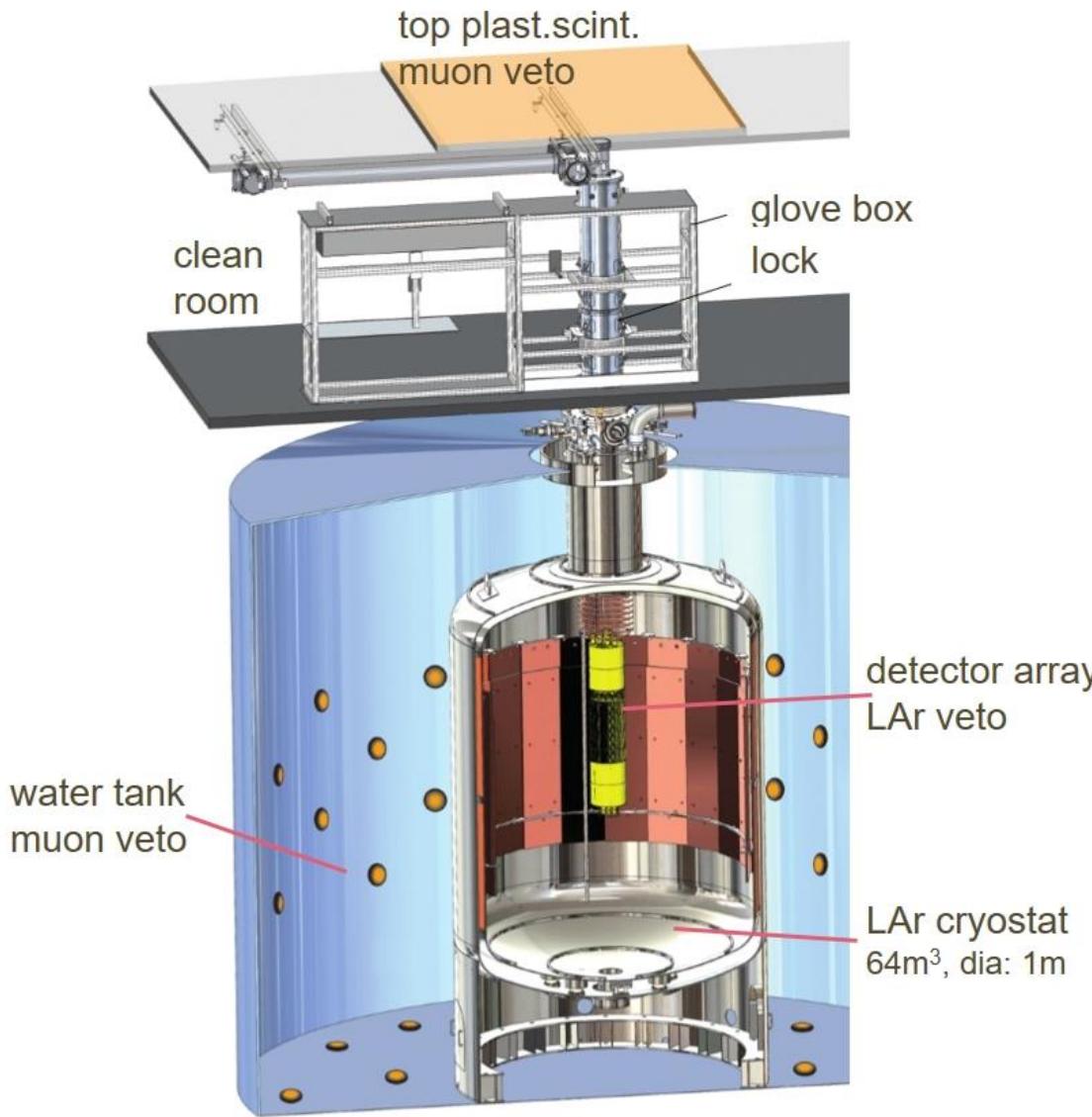
$$T_{1/2}^{0\nu\beta\beta} > 2.7 \cdot 10^{25} \text{ yr}$$

background in ROI
18 / FWHM· t_{isotope} ·yr



GERDA

$\Delta L \neq 0$



$\Delta E \sim 3\text{keV FWHM}$

LAr veto
combines

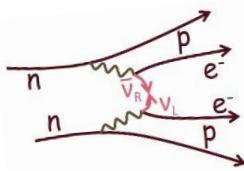
selfshielding / veto of liquid noble gas
high resolution of Ge detectors

pulse shape discrimination
multi site and surface event recognition

GERDA Phase II
started Dec 2015
35,6 kg enr. Ge (86%)

goals:

- 100 kgyr
- sensitivity > 10^{26} yr
- bkg 3 / FWHM t yr

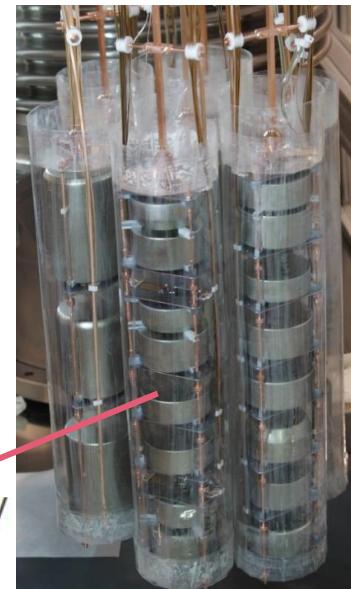
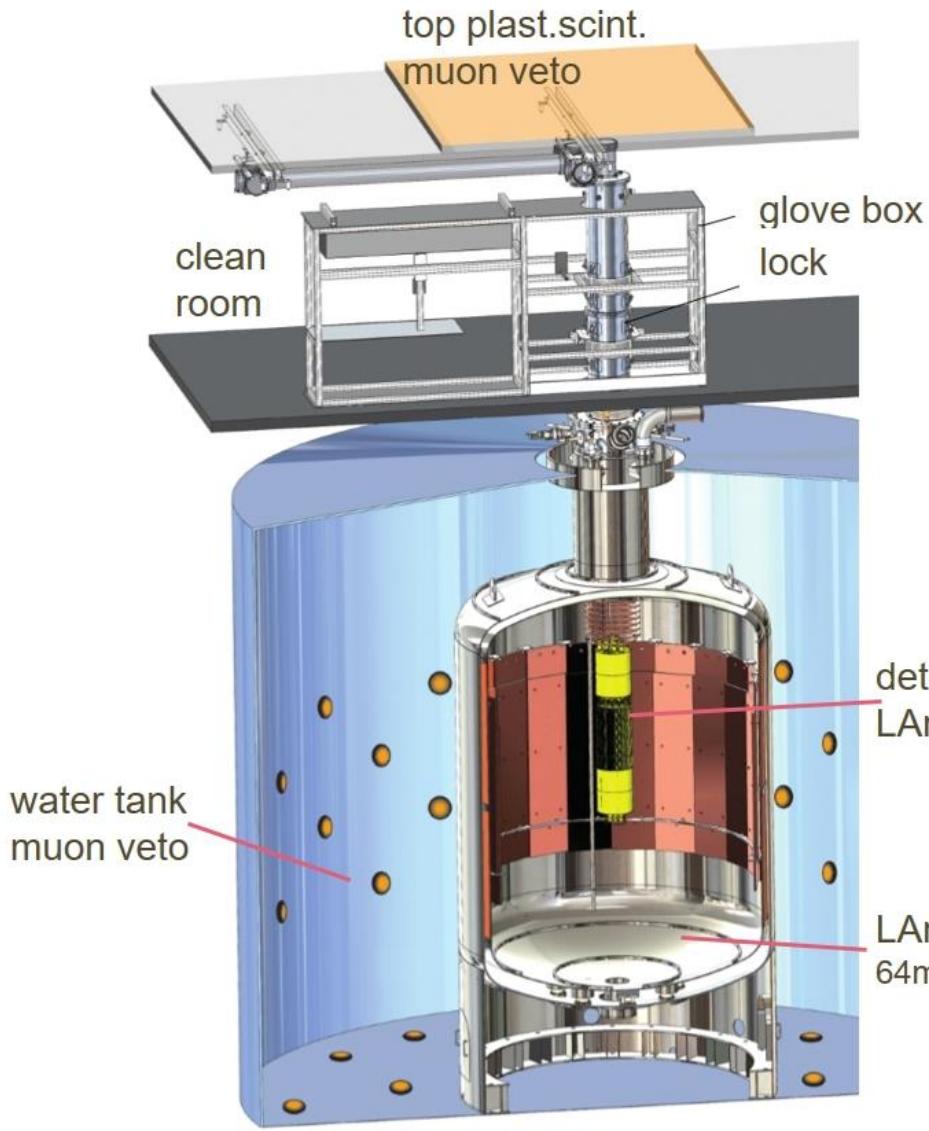


GERDA

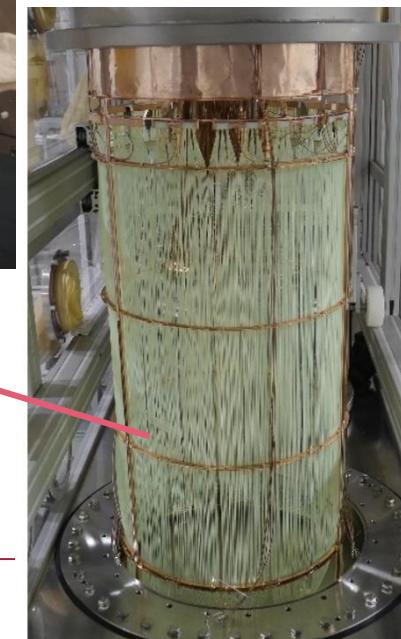
$\Delta L \neq 0$

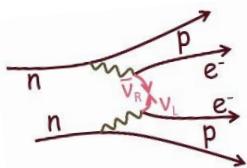
combines advantages

- active shielding of liquid noble gas
- high energy resolution of Ge detectors



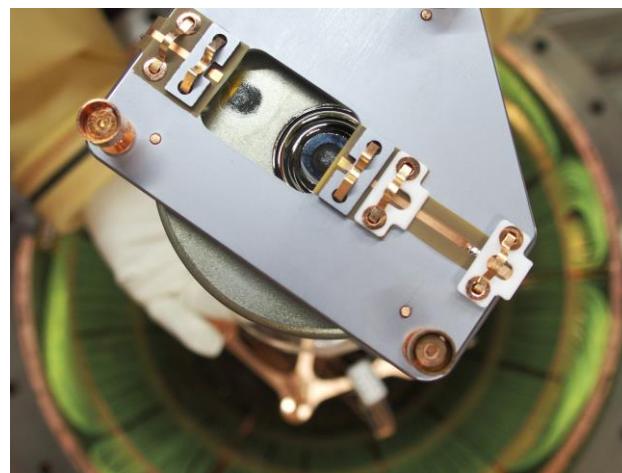
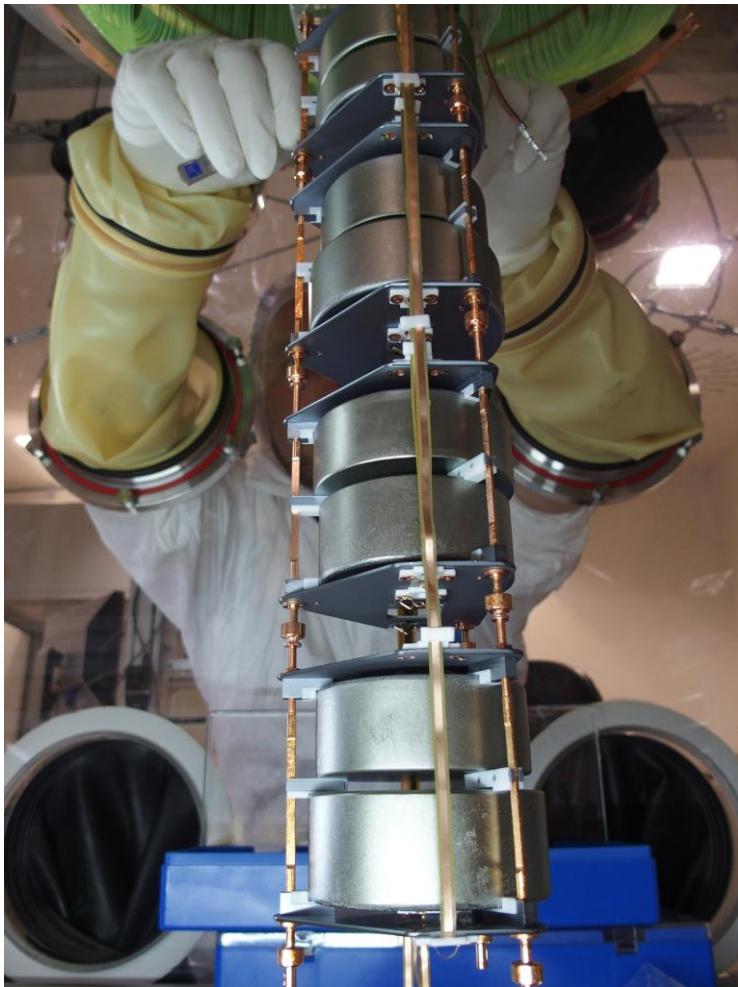
Gran Sasso

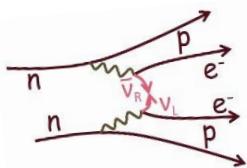




GERDA

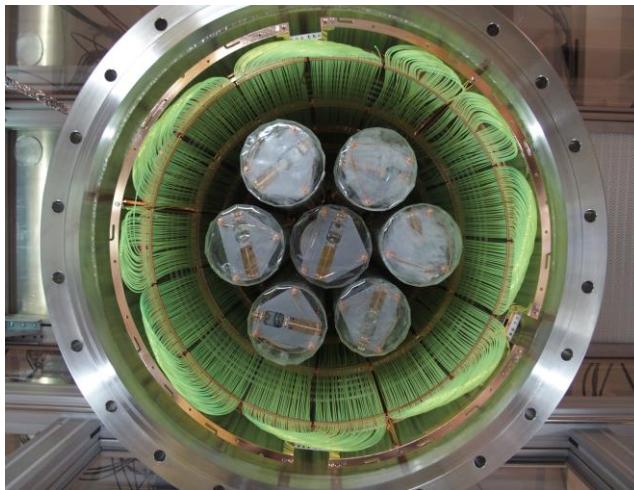
$\Delta L \neq 0$

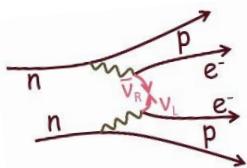




GERDA

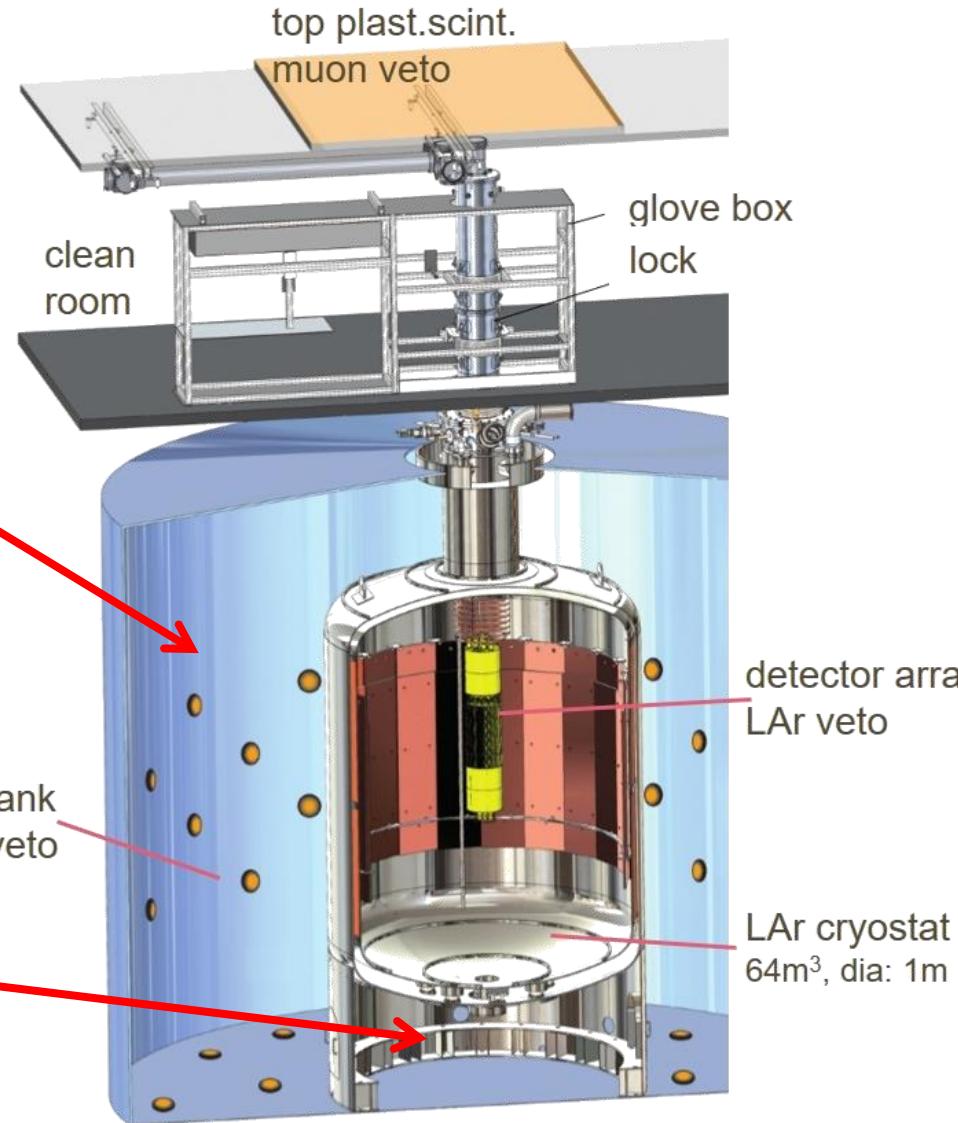
$\Delta L \neq 0$

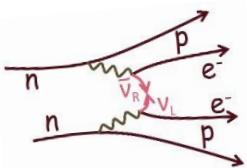




GERDA

$\Delta L \neq 0$

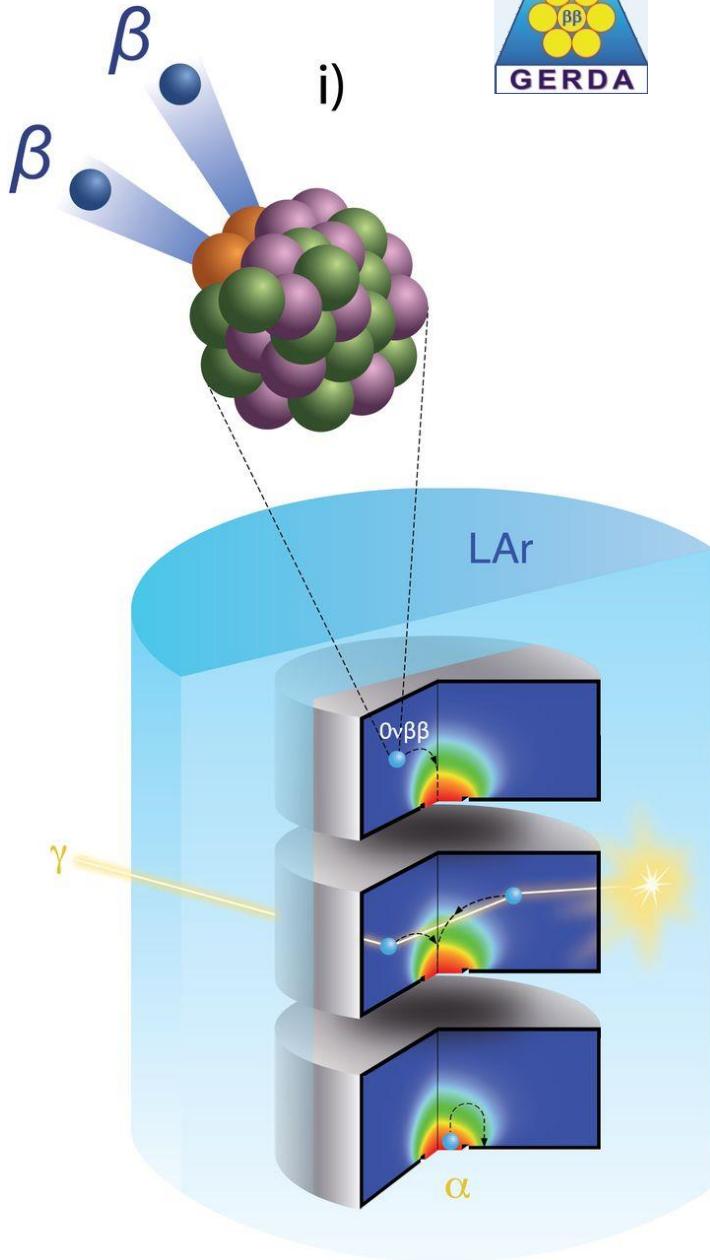




GERDA

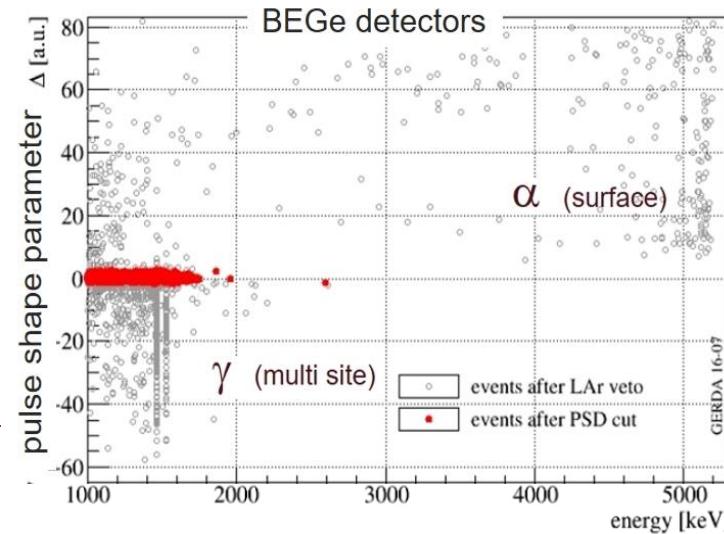
$\Delta L \neq 0$

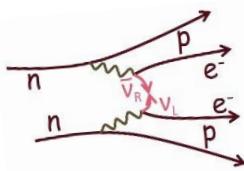




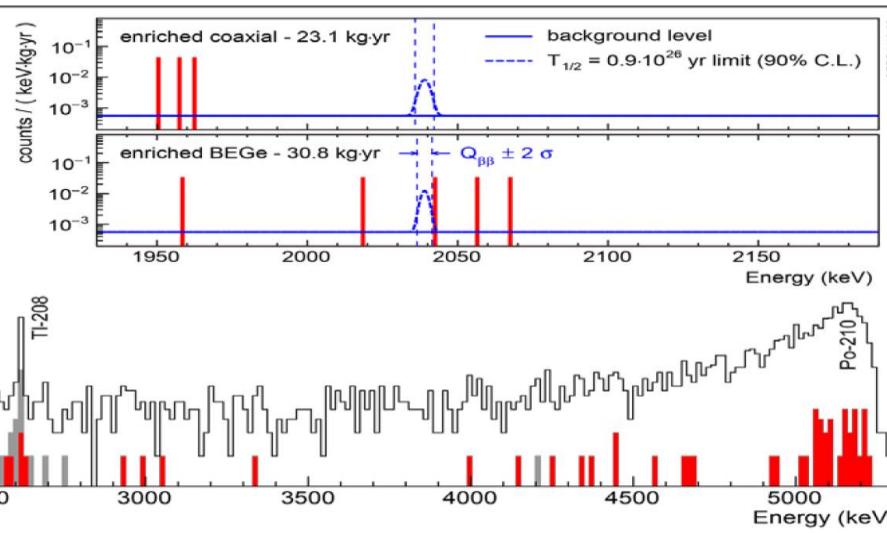
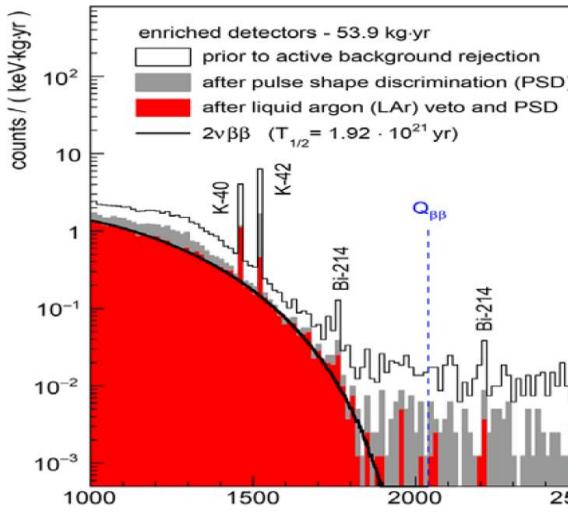
pulse shape discrimination multi site and surface event recognition

- i) $\beta\beta$ decay is localized (pointlike) – single site event
- ii) signal generation mostly when charges reach region close to small electrode
- iii) different drift times for multi site events
=> slower signal rise
- iv) α decays at unprotected contact surface show very fast rise

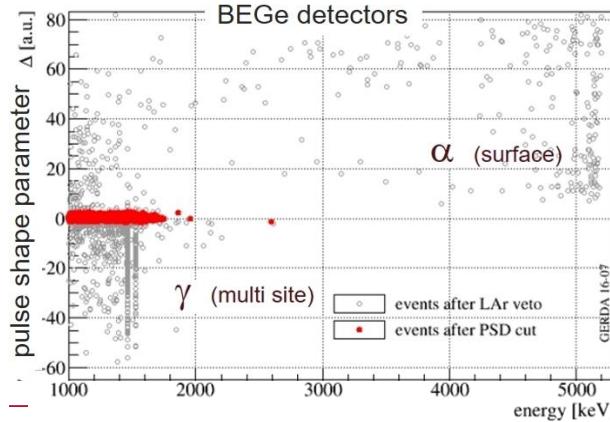




$\Delta L \neq 0$

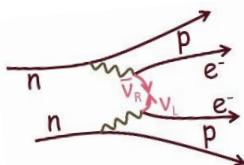


- 97% of events between 600-1300keV are $2\nu\beta\beta$
- Background 250 times lower compared to Heidelberg-Moscow Exp. (~10y)



pulse shape discrimination

- single site / multi site
⇒ γ lines suppressed by ~ 6
- surface / bulk
⇒ all α (surface) events removed



GERDA

$\Delta L \neq 0$

LAr veto +

pulse shape discrimination
multi site and surface event recognition

GERDA results

82.4 kg·yr total exposure

background in ROI

$4 / \text{FWHM} \cdot t_{\text{isotope}} \cdot \text{yr}$

sensitivity $11 \cdot 10^{25} \text{ yr}$

$T_{1/2}^{0\nu\beta\beta} > 9 \cdot 10^{25} \text{ yr}$

background free $0\nu\beta\beta$ experiment

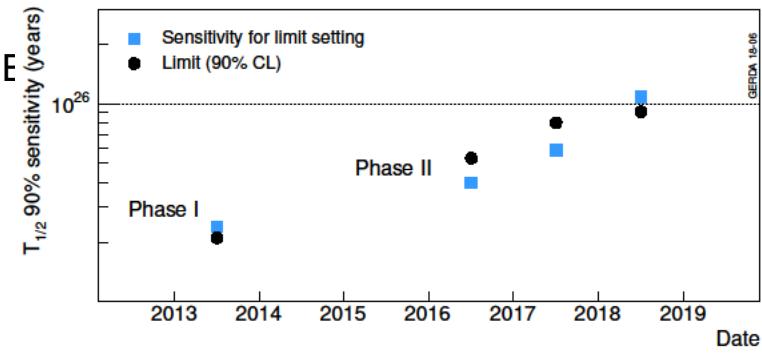
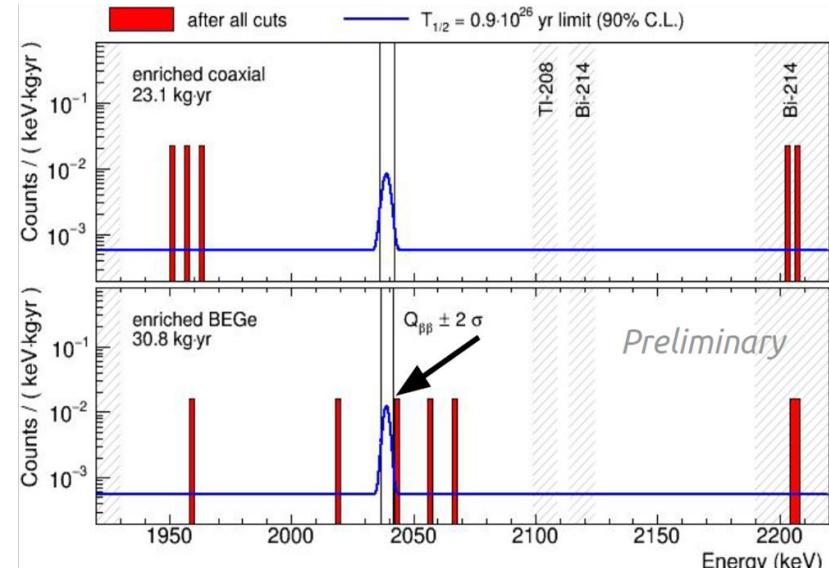
→ potential for discovery (up to $\sim 10^{26} \text{ yr}$)

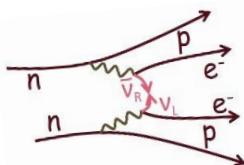
background 250 times lower compared to Heidelberg-Moscow E

makes sense to grow larger

(background goal for LEGEND 200 almost reached)

Science 10.112 / science.aav8613 (2019)

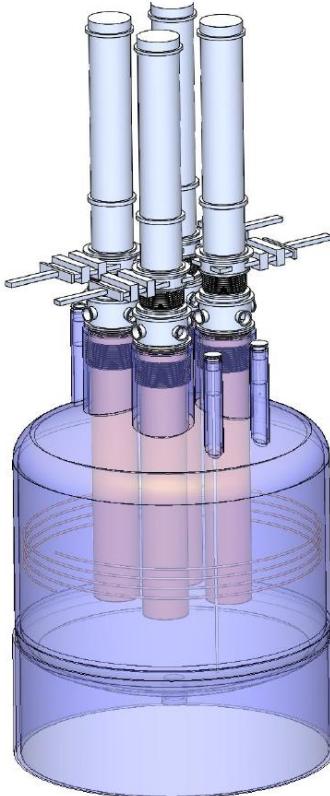




Search for Neutrinoless Double Beta Decay

$\Delta L \neq 0$

new collaboration formed LEGEND
Majorana + GERDA members + others



use GERDA concept and
staged approach to 1000kg



⇒ one worldwide collaboration on ^{76}Ge

LEGEND 200: first 200kg in GERDA setup @ Gran Sasso

- starting 2021
- ^{76}Ge available for 190kg of detectors
- funded by NSF, INFN, MPI, BMBF

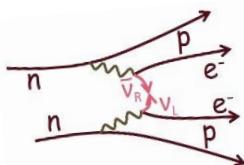
sensitivity $> 10^{27} \text{ yr}$



Large Enriched
Germanium Experiment
for Neutrinoless $\beta\beta$ Decay

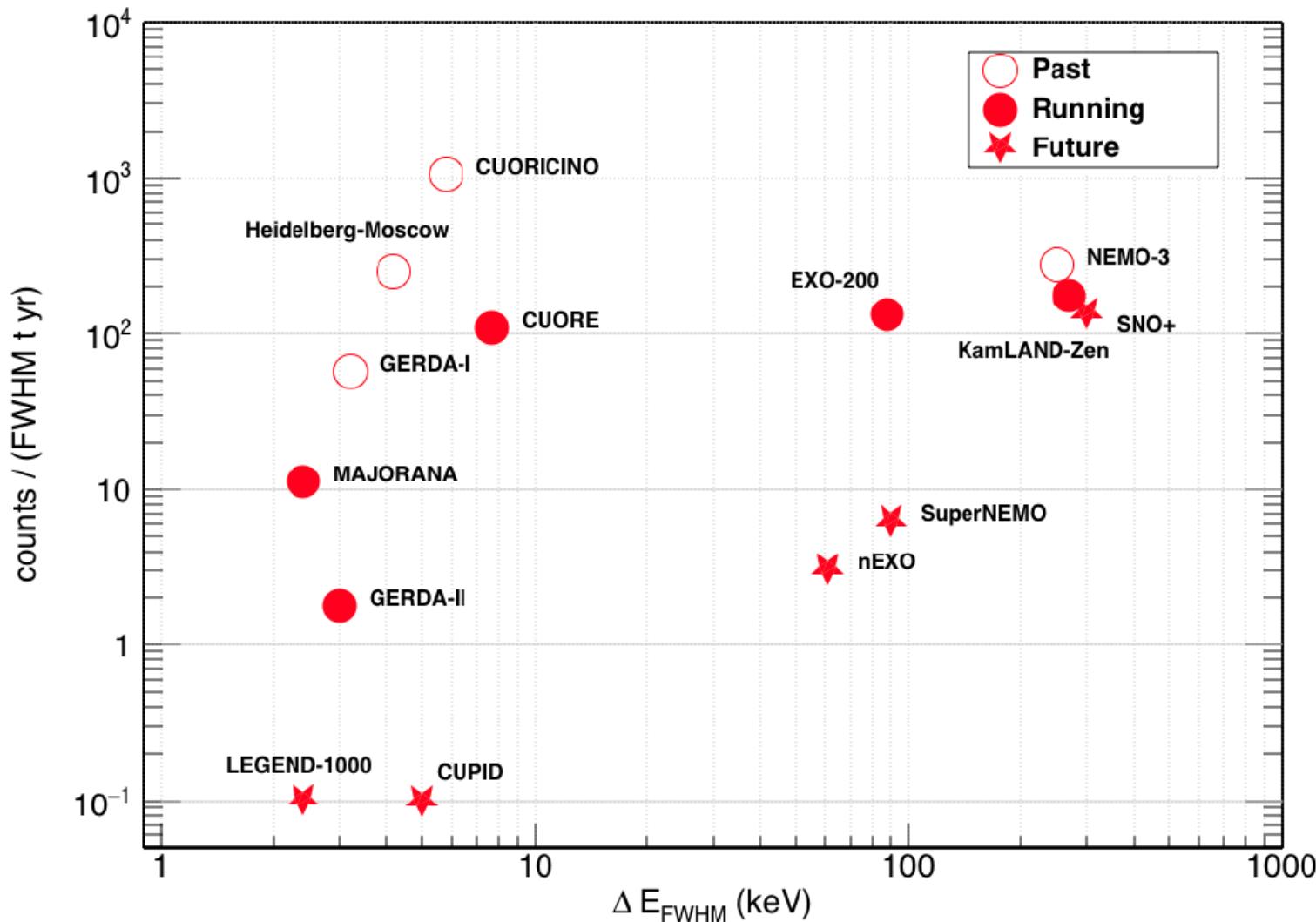
sensitivity $> 10^{28} \text{ yr}$

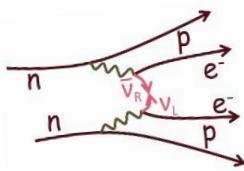
LEGEND 1000: 1000kg phase depends on US down selection process



Search for Neutrinoless Double Beta Decay

$\Delta L \neq 0$

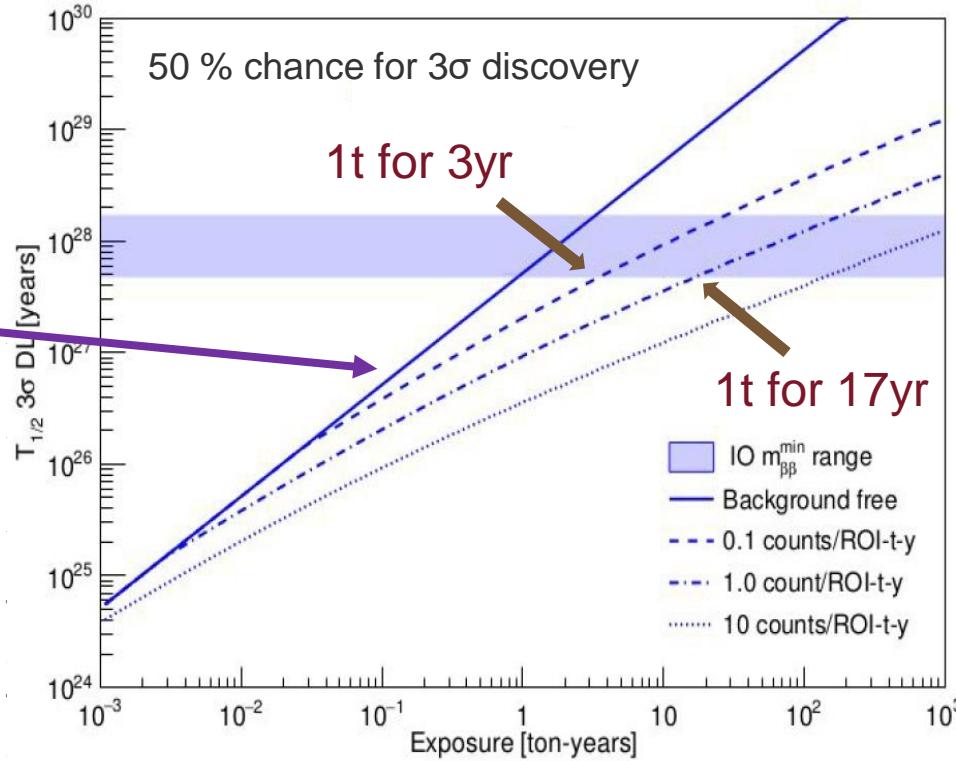




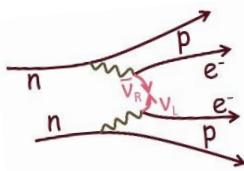
Search for Neutrinoless Double Beta Decay

$\Delta L \neq 0$

background
free regime
 $T_{1/2} \sim \text{exposure}$

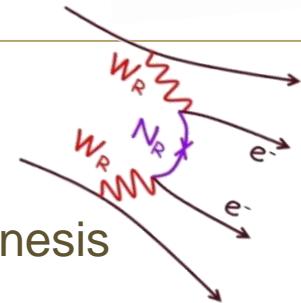


low background essential for discovery potential



Summary

$\Delta L \neq 0$



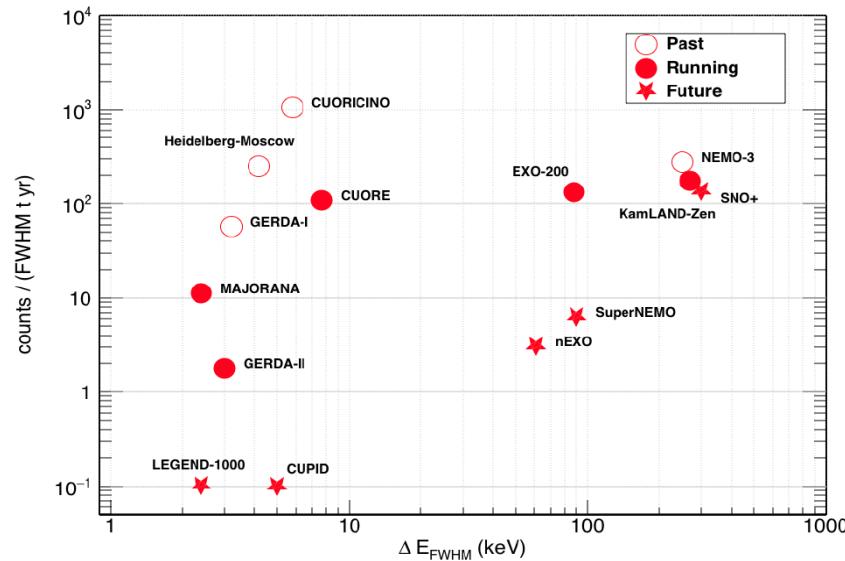
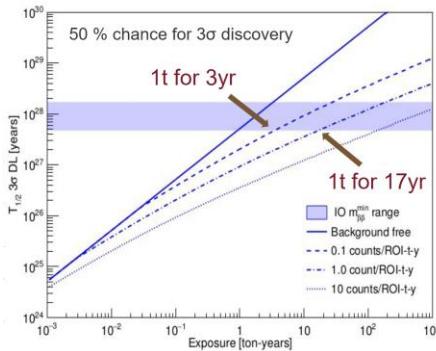
- search for double beta decay highly motivated:

$\Delta L \neq 0$, Majorana ν , lightness of ν -mass, Leptogenesis

next experiments explore range up to $T_{1/2} < 10^{27}$ yr mid term

\Rightarrow chance for discovery of $\Delta L \neq 0$ $T_{1/2} < 10^{28}$ yr long term

- field is very active and competitive,
variety of approaches and technologies



1902.04097

low background essential for discovery potential

END
