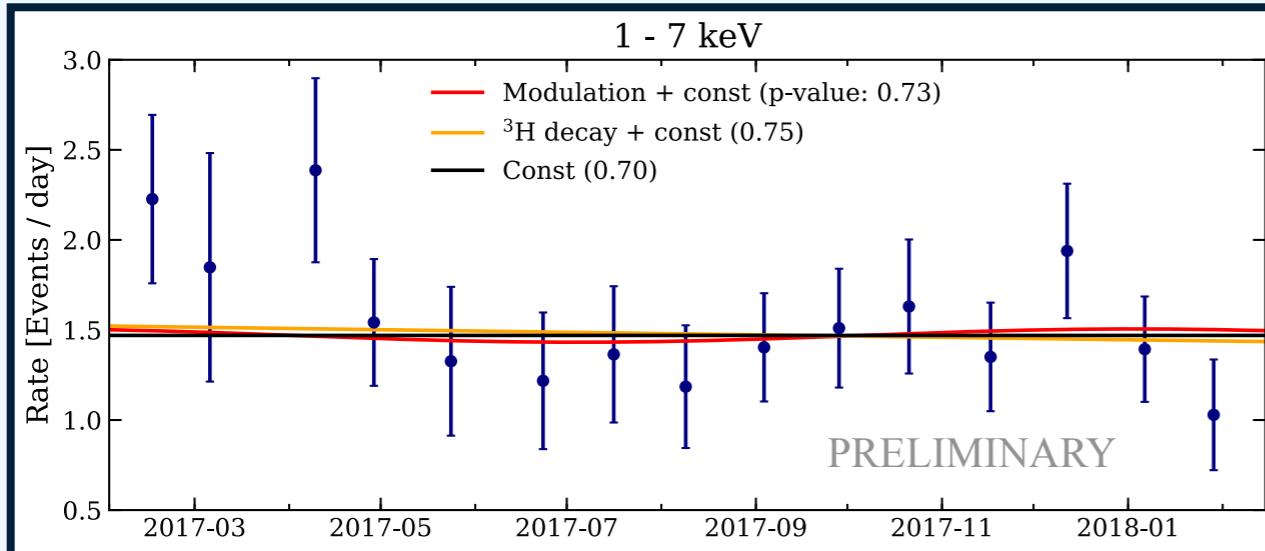
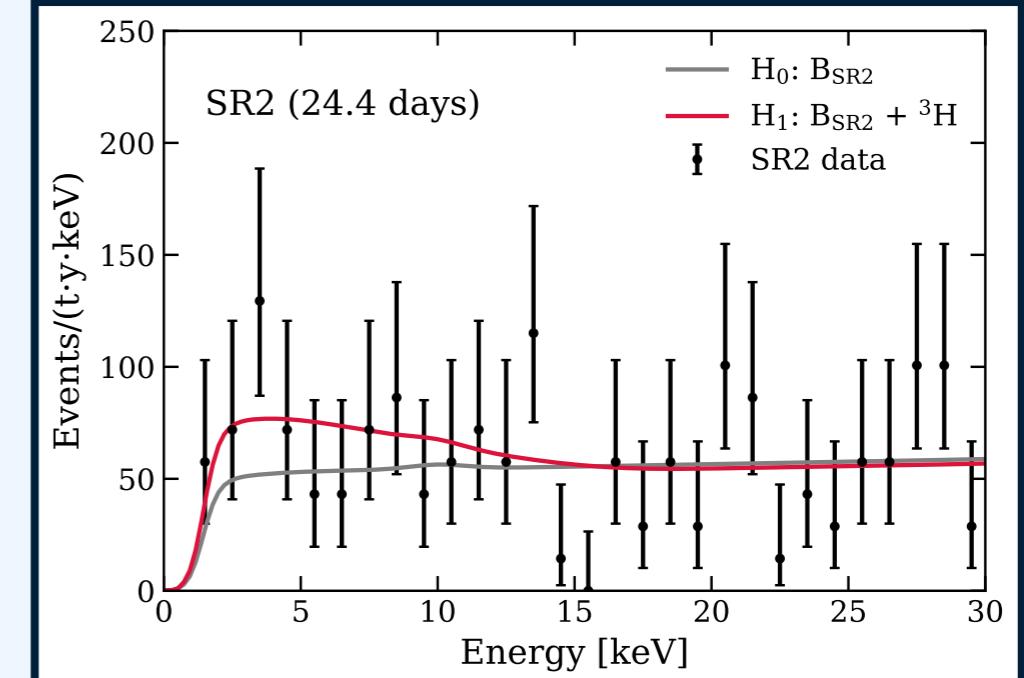


Additional checks

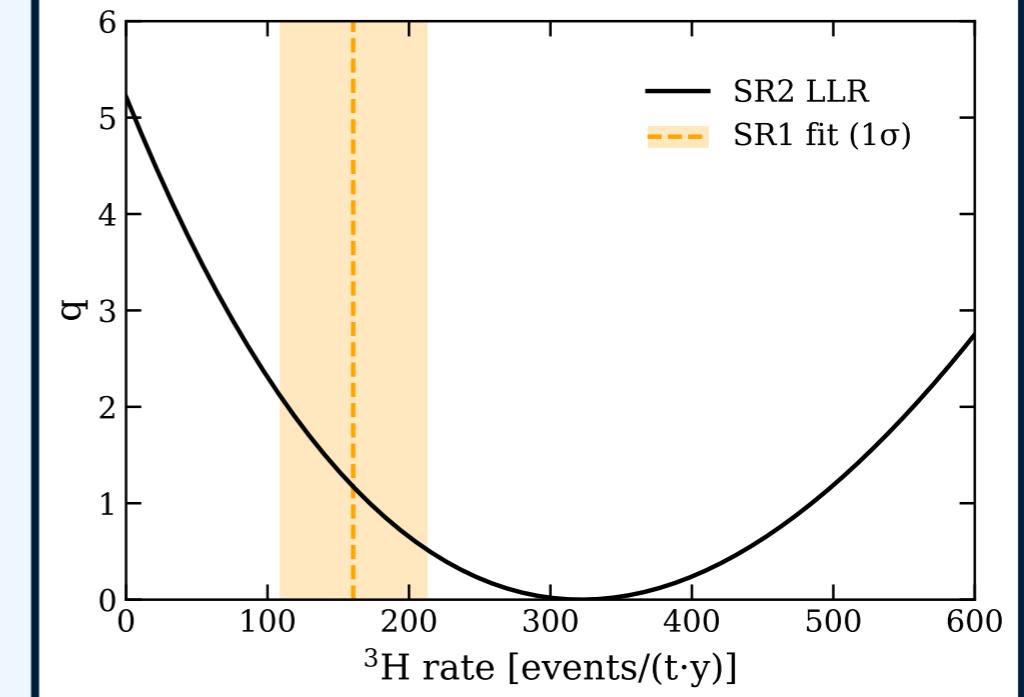
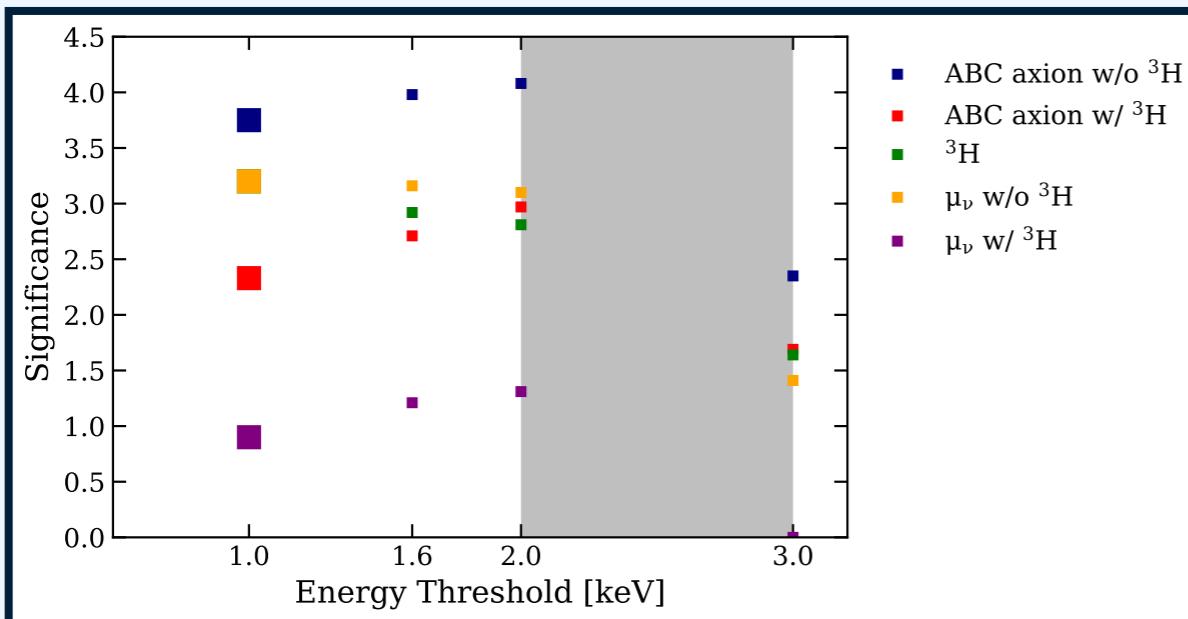
Time dependence



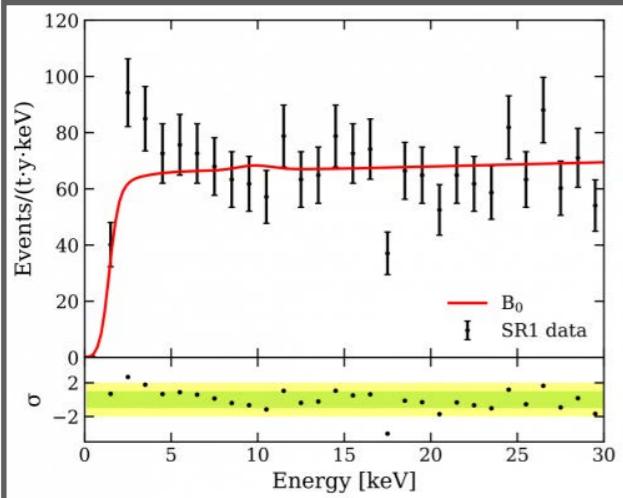
Additional data



Threshold effects



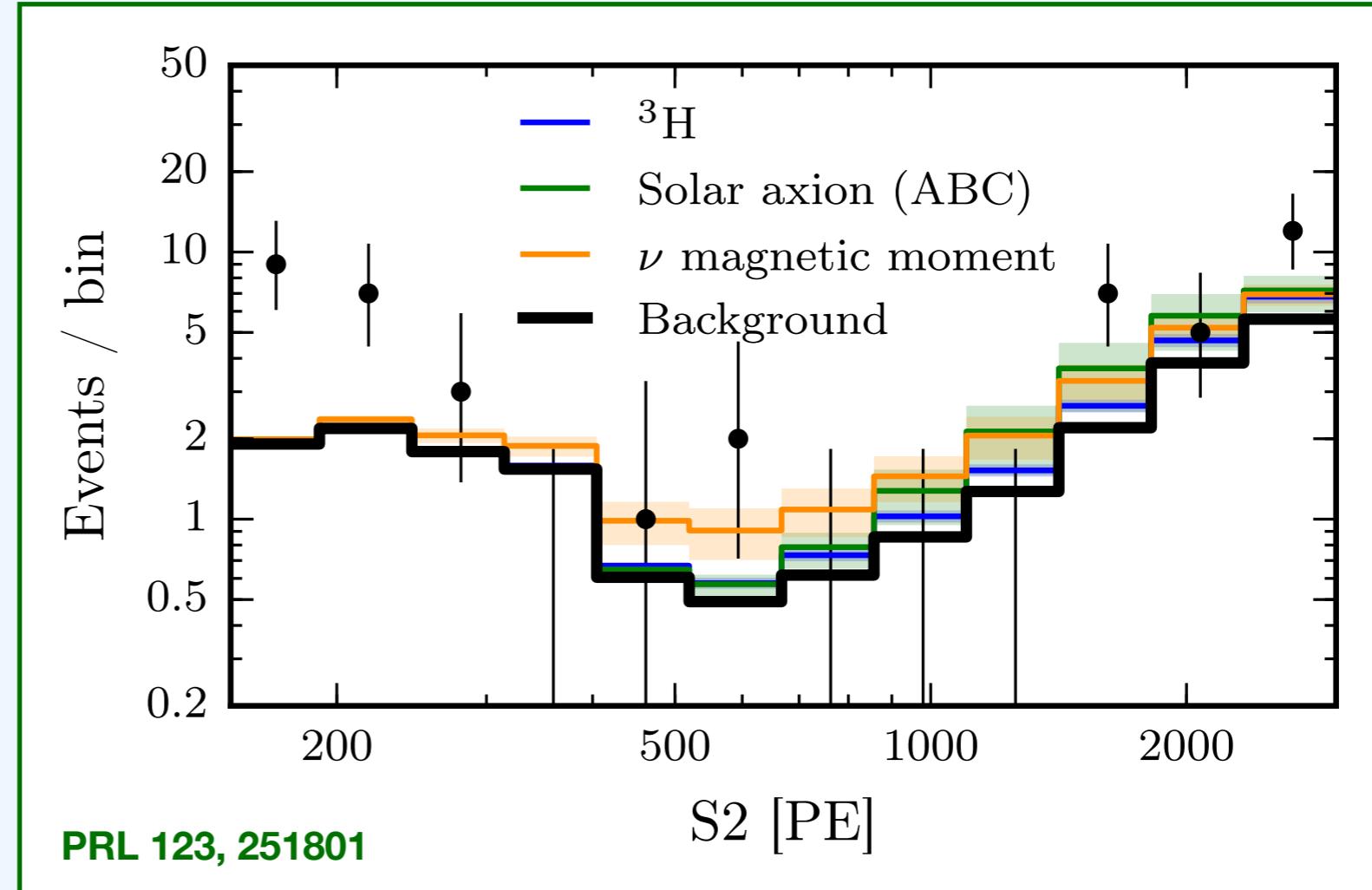
S₂-only analysis



S₂-only analysis

limit-setting only

allows for a lower
energy threshold

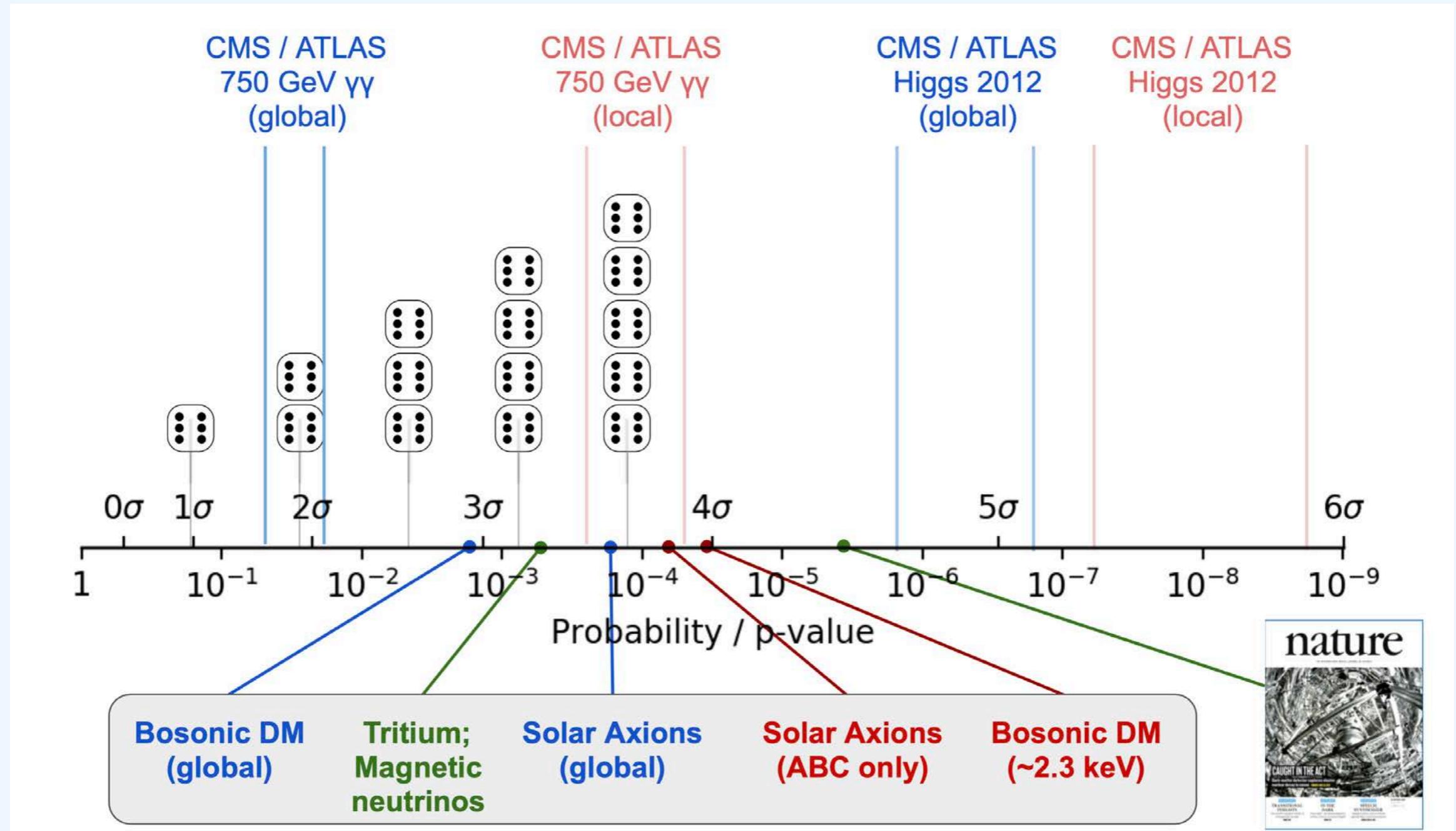


$$\begin{aligned}\mu_\nu &< 3.1 \times 10^{-11} \mu_B \\ g_{ae} &< 4.8 \times 10^{-12} \\ R_{^3\text{H}} &< 2256 \text{ events/t/y}\end{aligned}$$

**consistent with this work
for all 3 hypotheses**

Putting it all together.

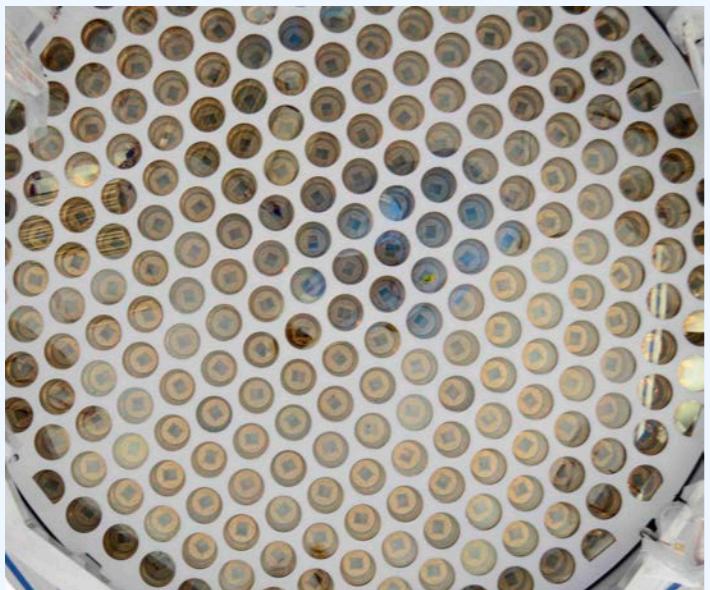
Multiple hypotheses



Our results are...
inconclusive.

(what's next?)

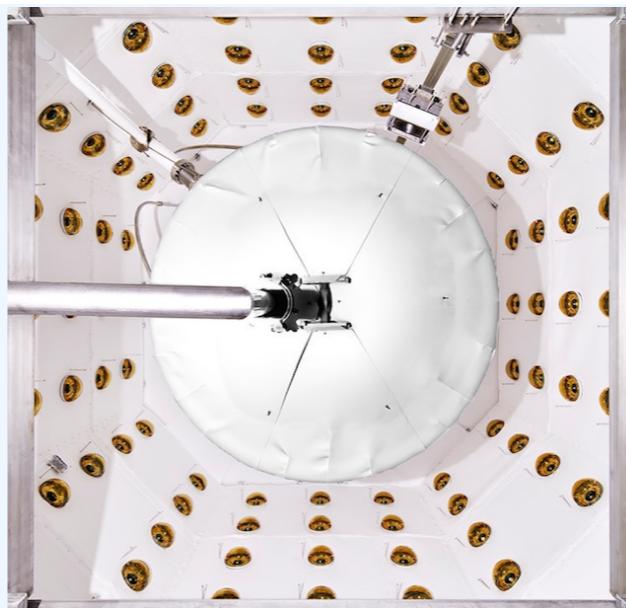
XENONnT



PMT array (494 PMTs in total,
in 2 arrays)



TPC (5.9 t LXe,
4 t fiducial)



Neutron veto (120 PMTs,
Gd-doped water)

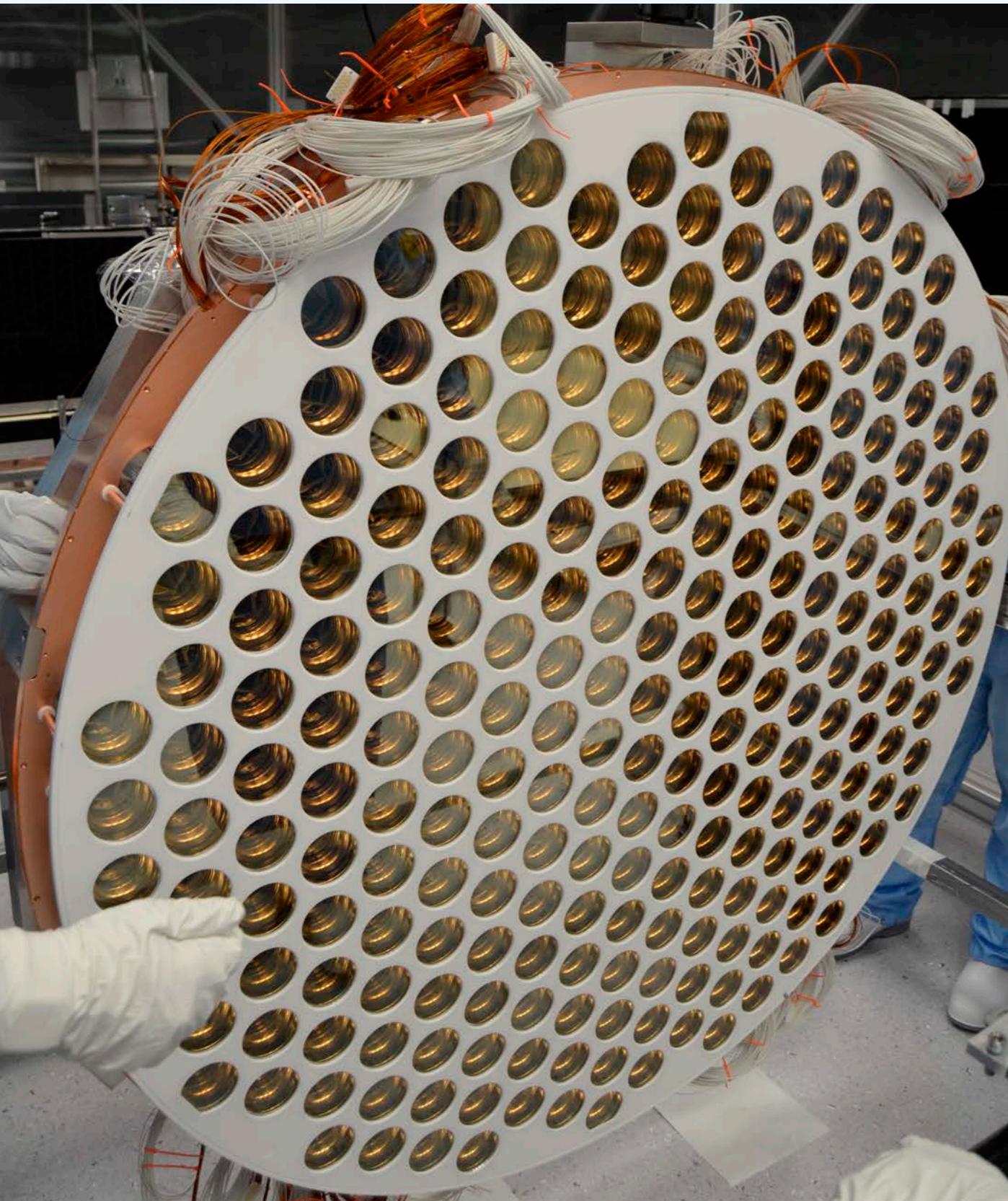


LXe purification system
(faster purification)

Rn distillation column
reduce ^{222}Rn (^{214}Pb)



XENONnT

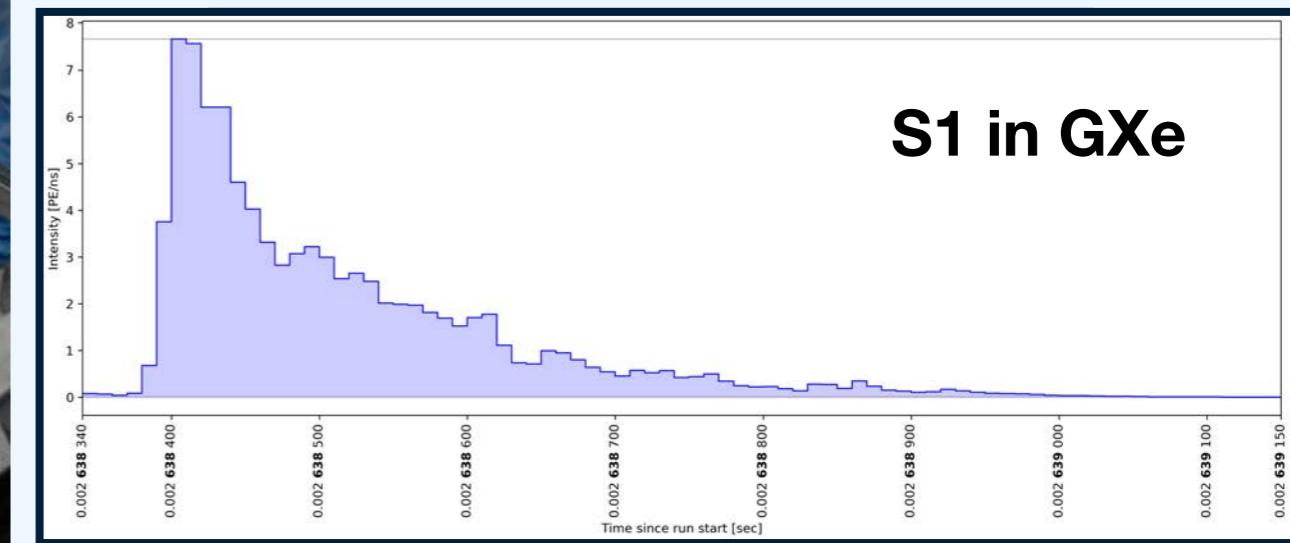


3x

Larger active volume

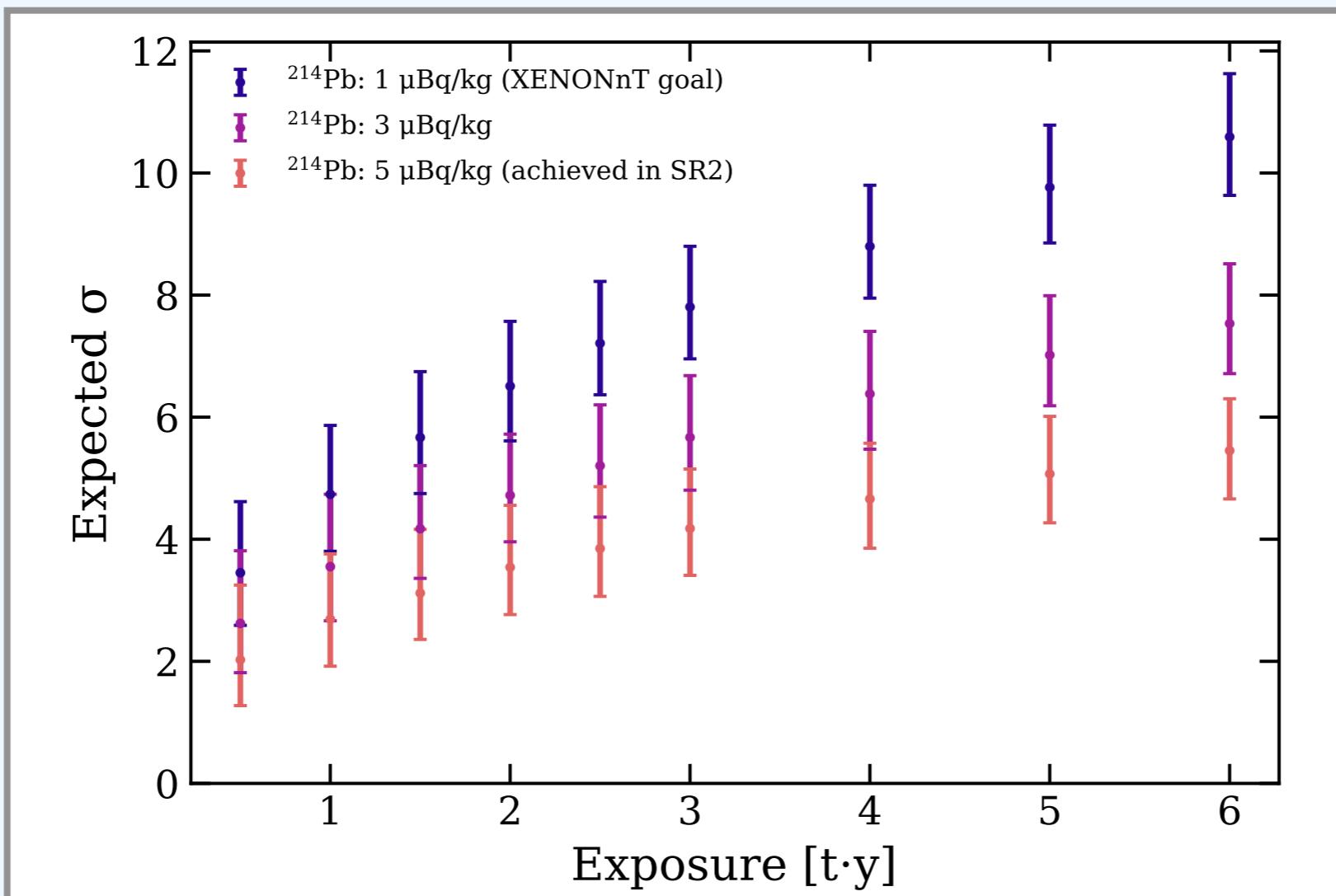
~1/6

Reduced background
level



S1 in GXe

XENONnT



Based on energy spectrum alone.

Uses best-fit from 1T search.

XENONnT will discriminate axions from tritium with ~ few months of data

Conclusions

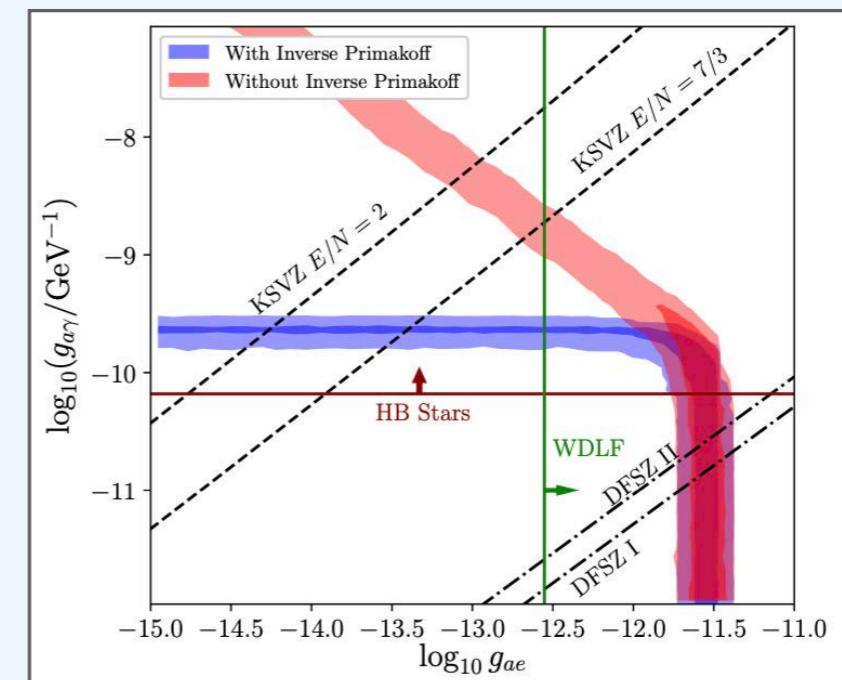
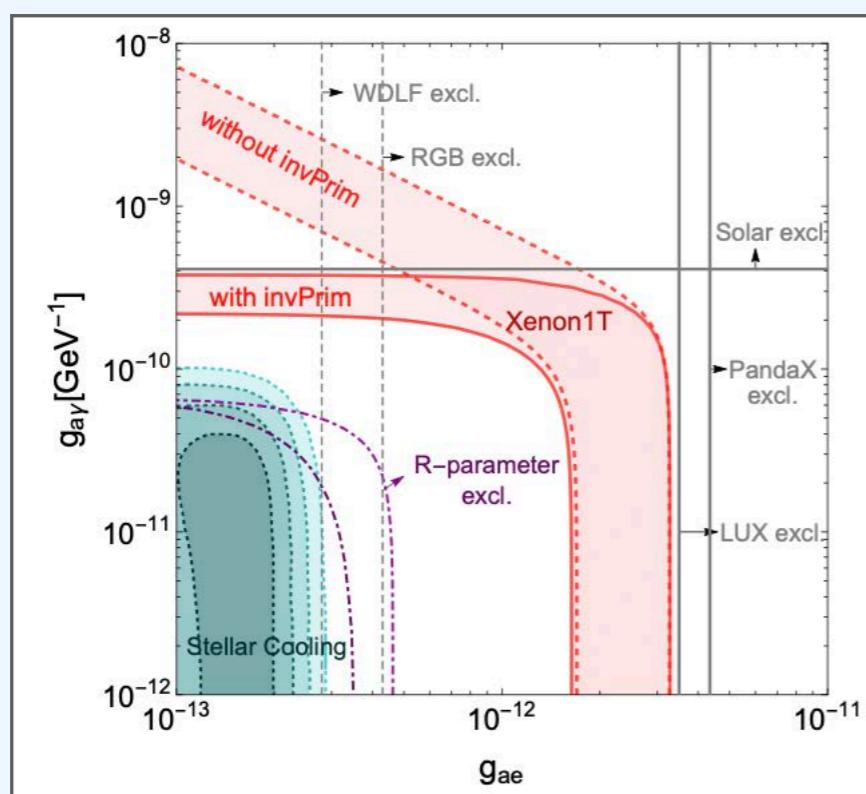
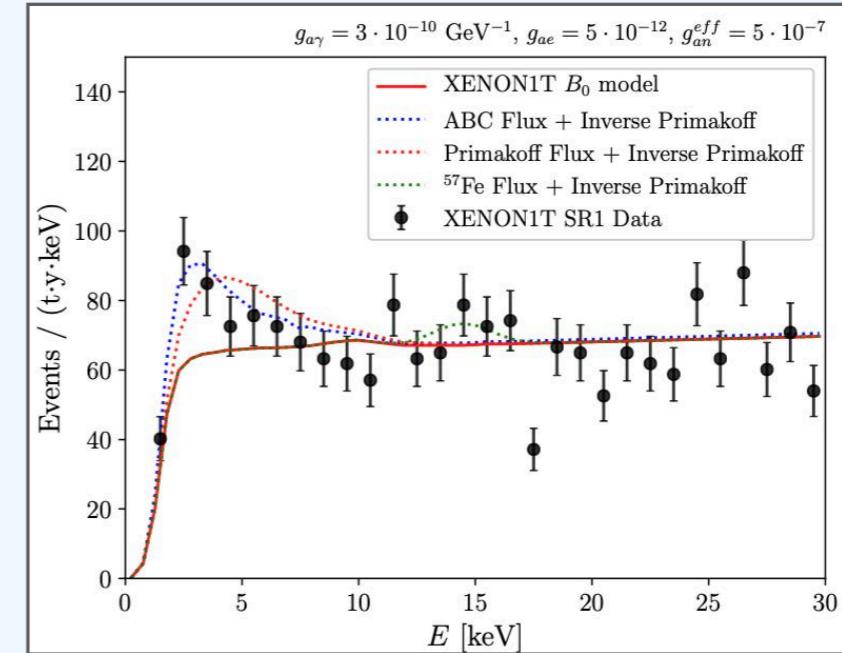
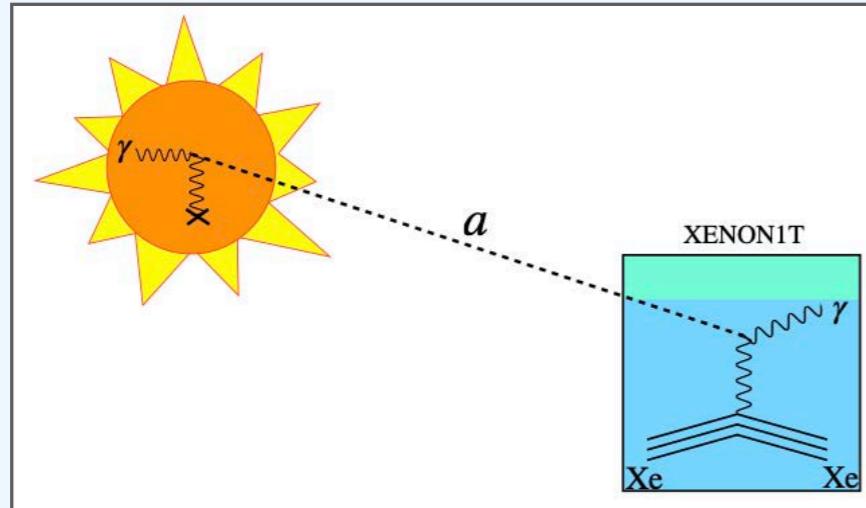
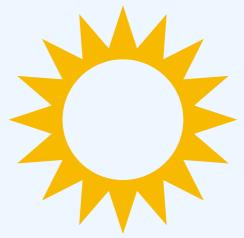
- Solar axions favoured over background at 3.4σ , *but*
- it is in strong tension with stellar constraints, *and*
- a tritium background at 3.2σ can neither be confirmed nor excluded.
- an ^{37}Ar -37 background is strongly constrained and not likely to explain the excess
- An enhanced neutrino magnetic moment favoured over background at 3.2σ (0.9σ w/ ^3H)
- Bosonic dark matter peak at $2.3 +/ - 0.2$ keV has 4.0σ local significance, but note 3.0σ global.

It is too soon to draw any conclusions; however

XENONnT is coming soon!



Inverse Primakoff Effect



C. Gao, et al.
arXiv:2006.14598

J.B. Dent, et al.
arXiv:2006.15118

Minimising the tension with stellar constraints

QCD axion models

$$m_a \simeq \frac{6 \times 10^6 \text{ GeV}}{f_a} \text{ eV/c}^2$$

DFSZ: two Higgs doublets model
couplings to leptons at tree level

Dine-Fischler-Srednicki-Zhitnitsky (DFSZ)

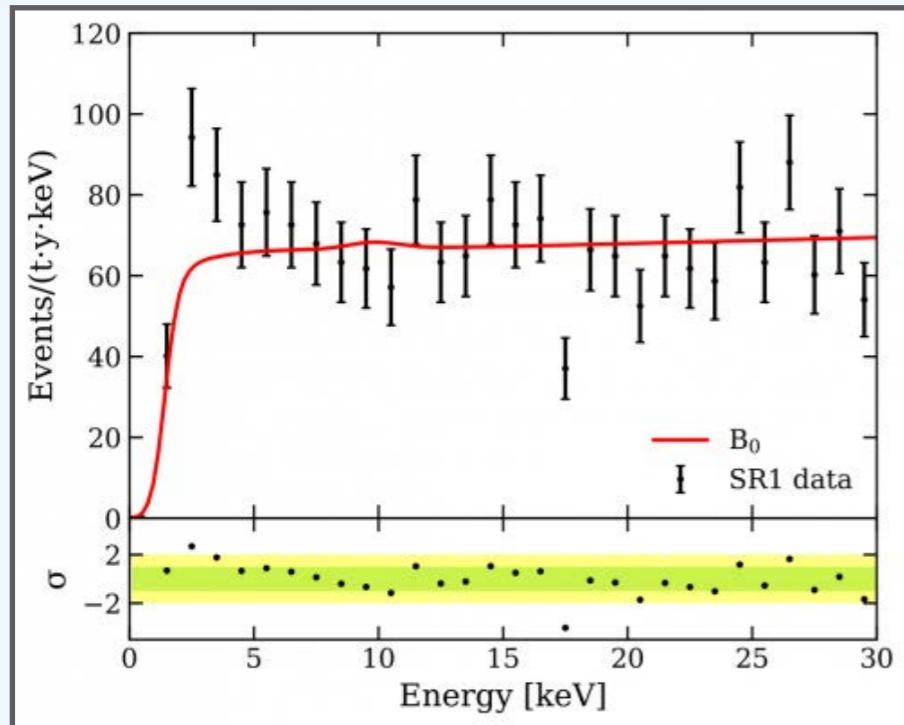
KSVZ: heavy quark model
couplings to leptons only at loop level

Kim-Shifman-Vainshtein- Zhakharov (KSVZ)

**axion-photon coupling
same for both models**

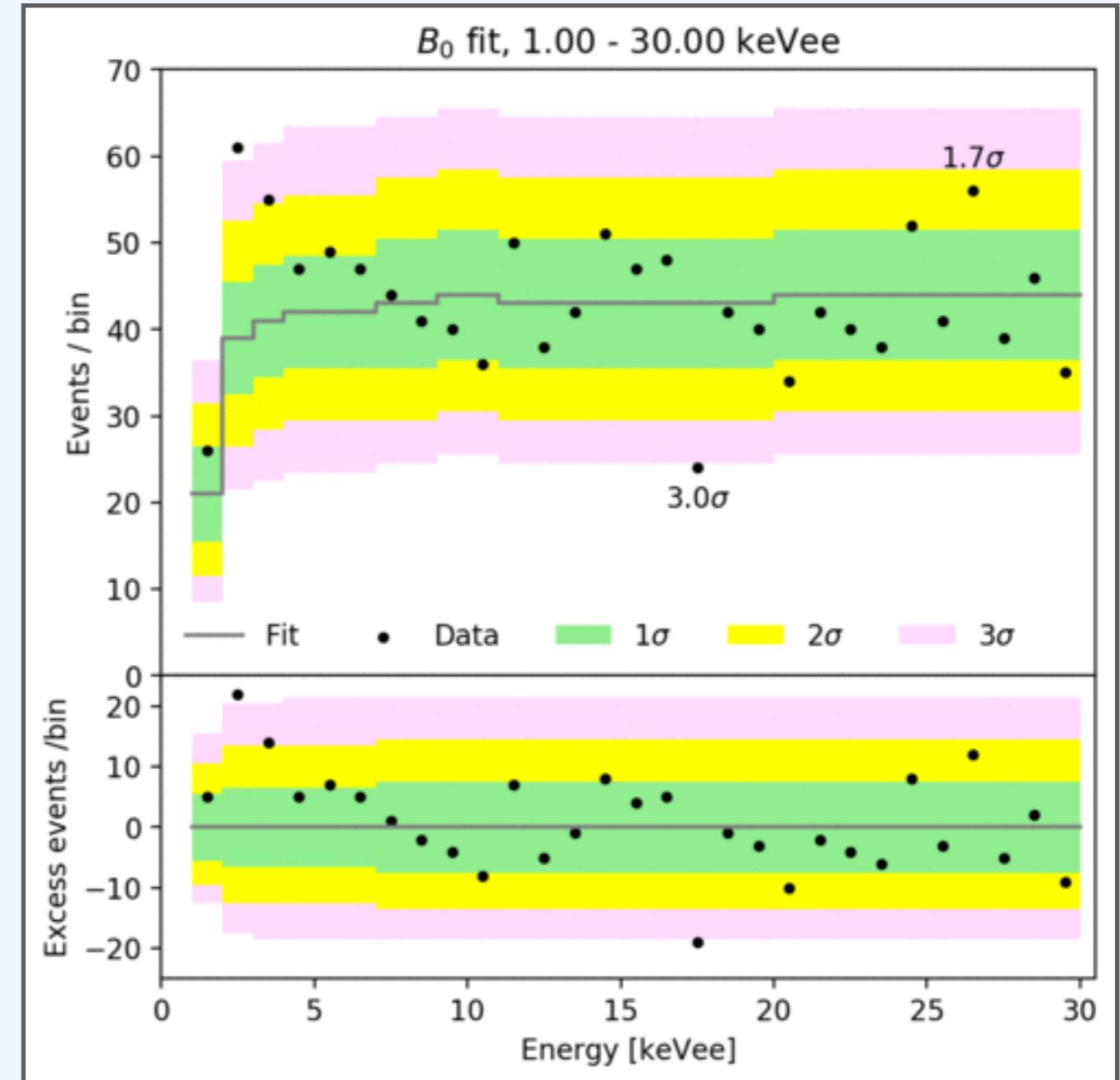
- relative contributions from each component can allow to distinguish between models (Primakoff dominates in KSVZ models); can also constrain β_{DFSZ}
- nuclear transition contribution always relatively small

Fluctuations and correlations



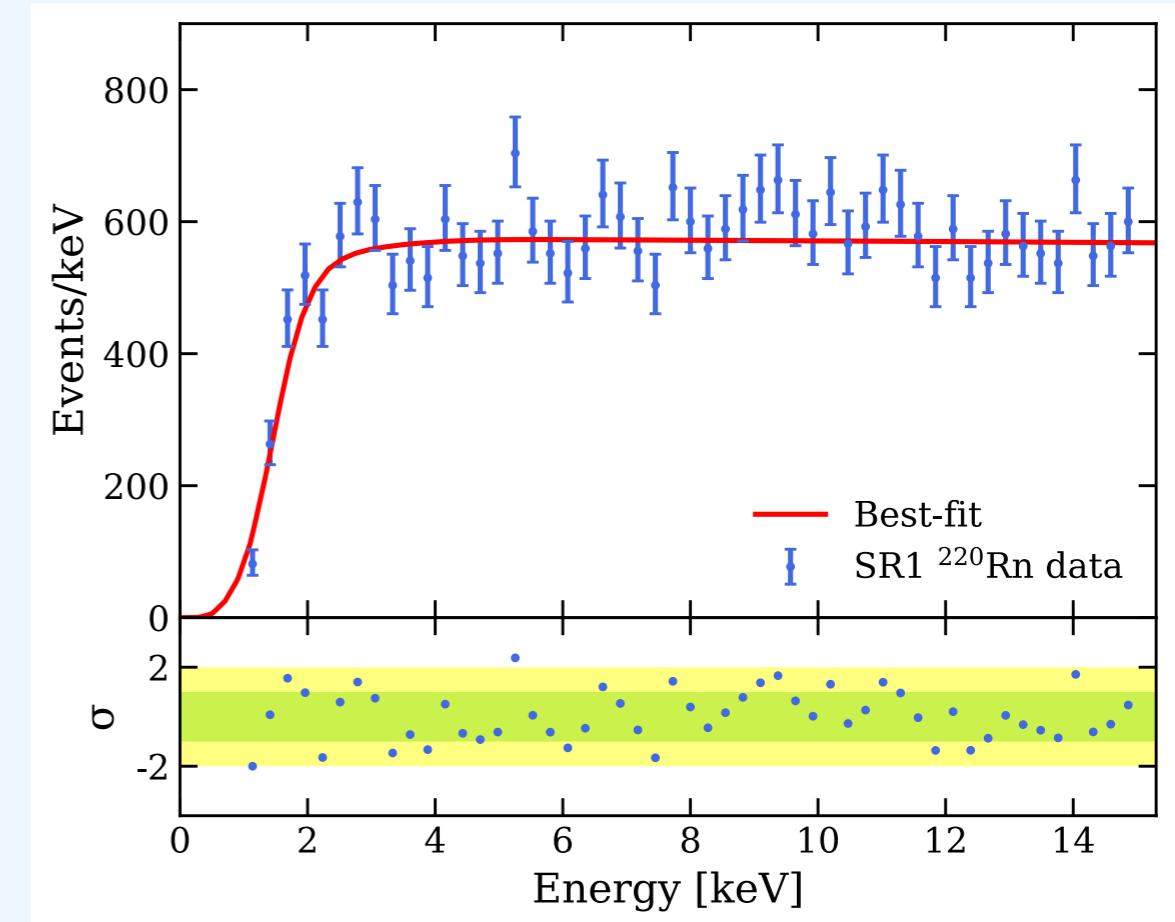
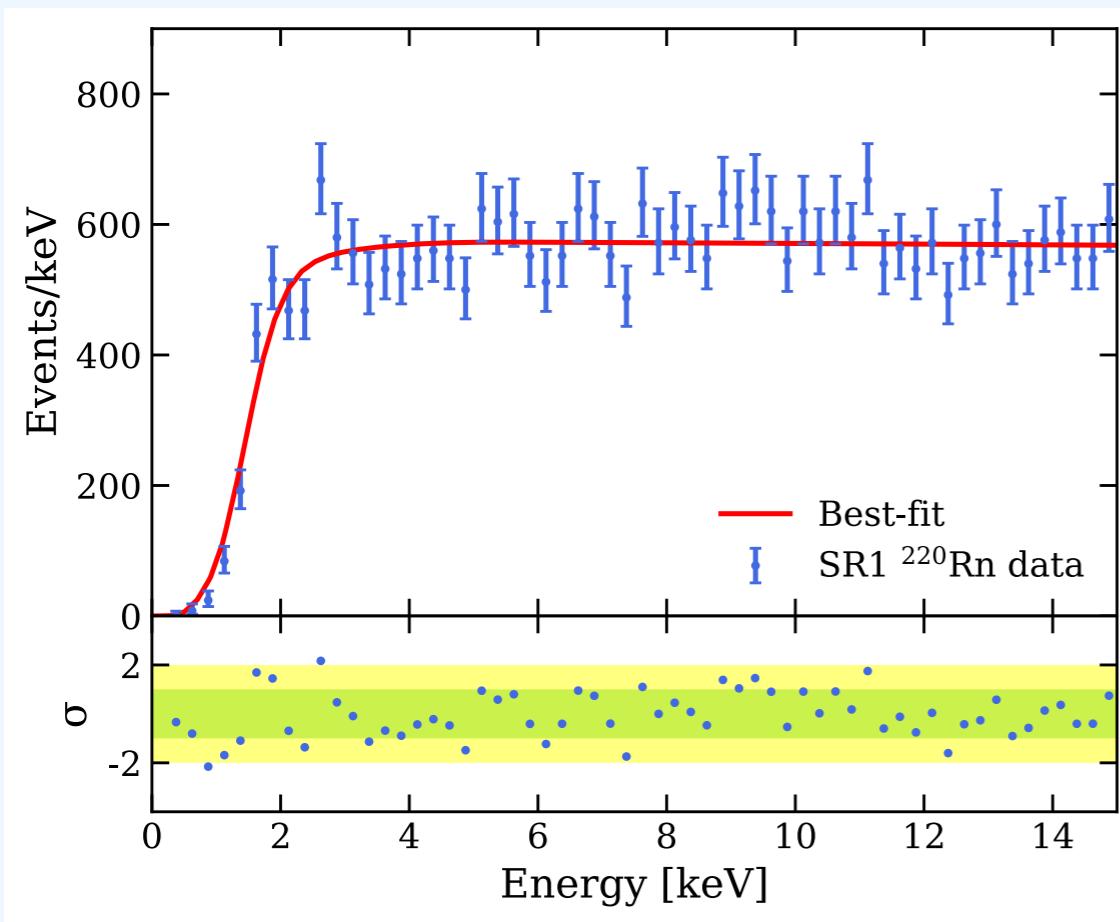
statistical fluke? (see 17 keV dip)

funny correlation?
(1-10 keV rising steadily)



Note: we use an unbinned profile likelihood analysis

Calibration data



^{220}Rn rebinned

SR2

Data taken after SR1 through end of 2018

Require strict data selection due to changing detector conditions

24.4 live-days in total used for cross-check

20% reduced background level

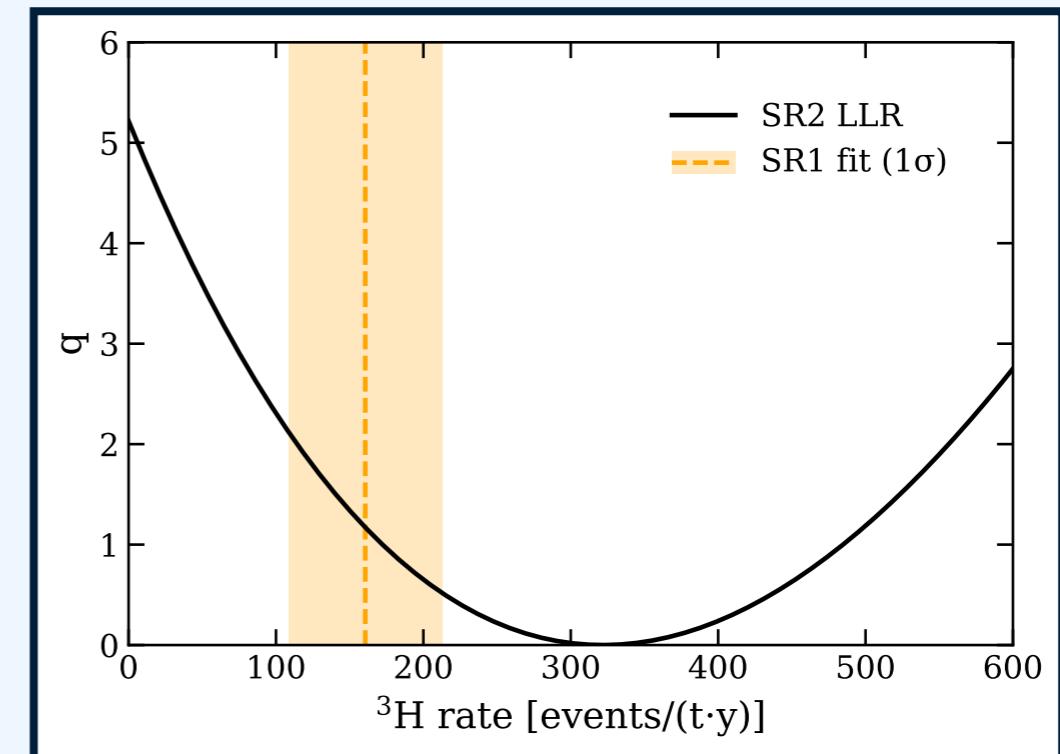
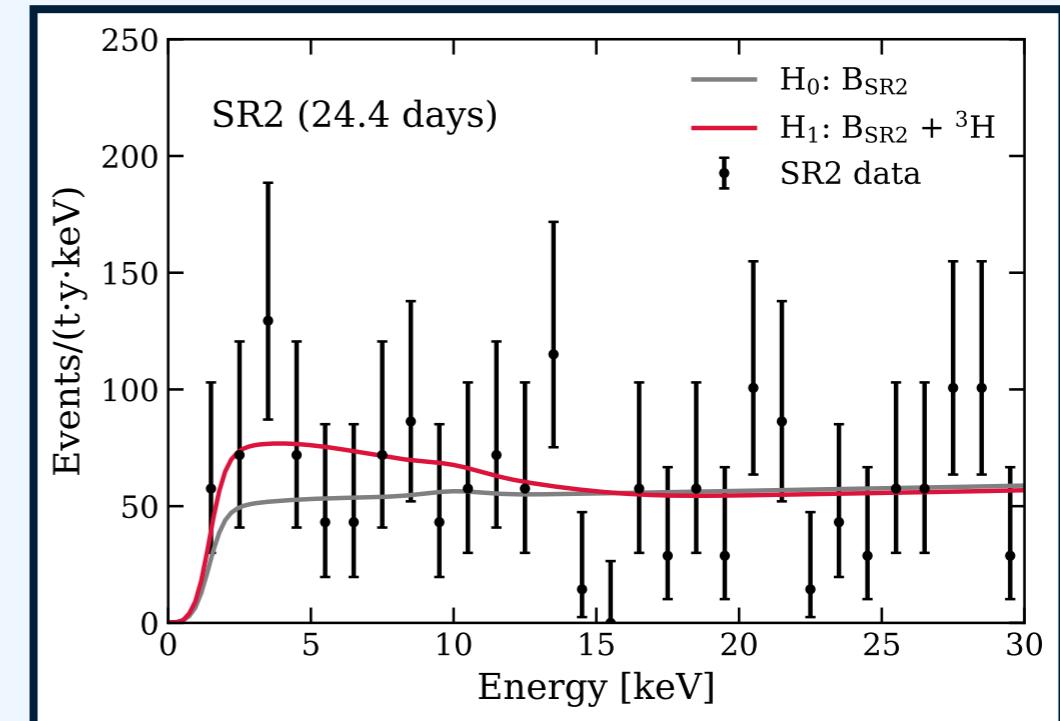
Improved purification

Goal: test the tritium hypothesis

Tritium favored over background-only at 2.3σ .

Rate consistent with SR1

Similar results for solar axion and magnetic moment ($\sim 2\sigma$)

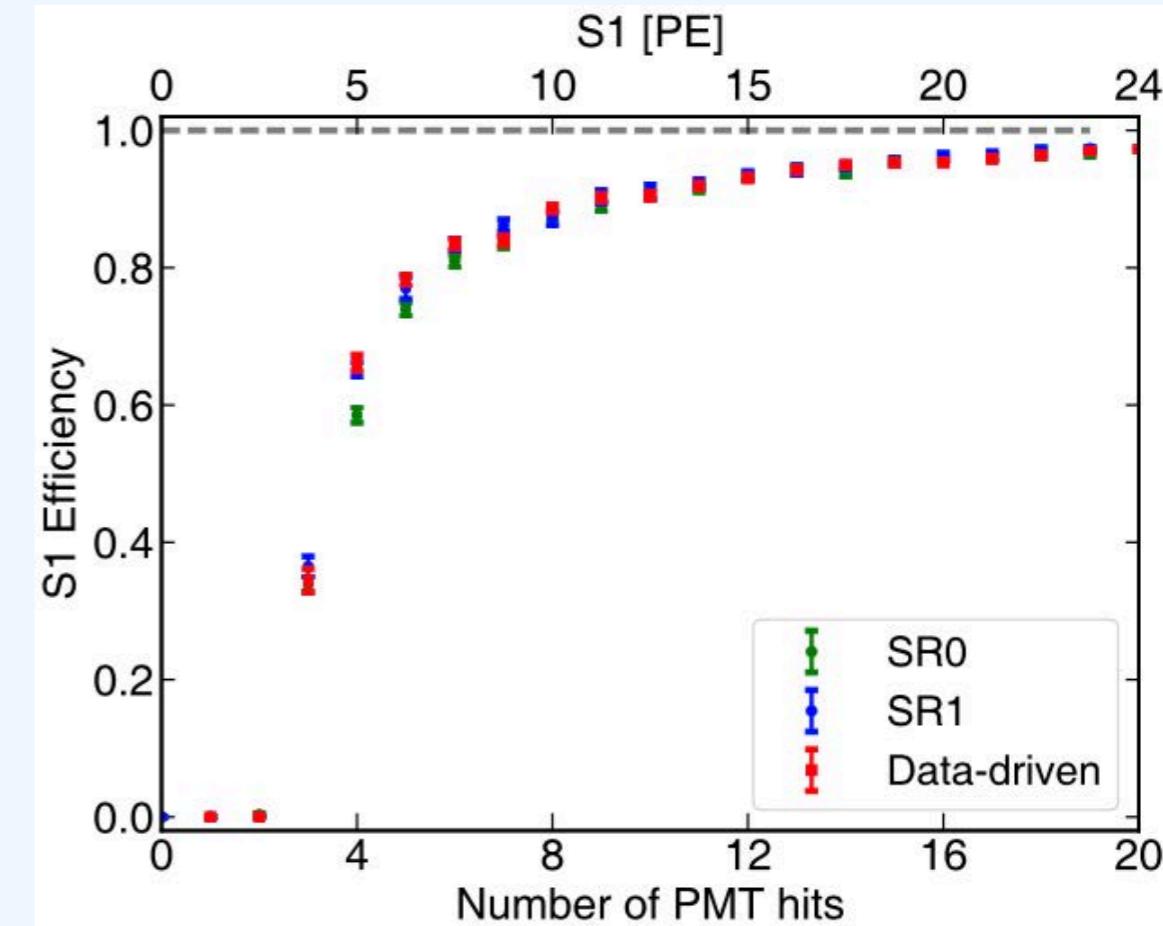
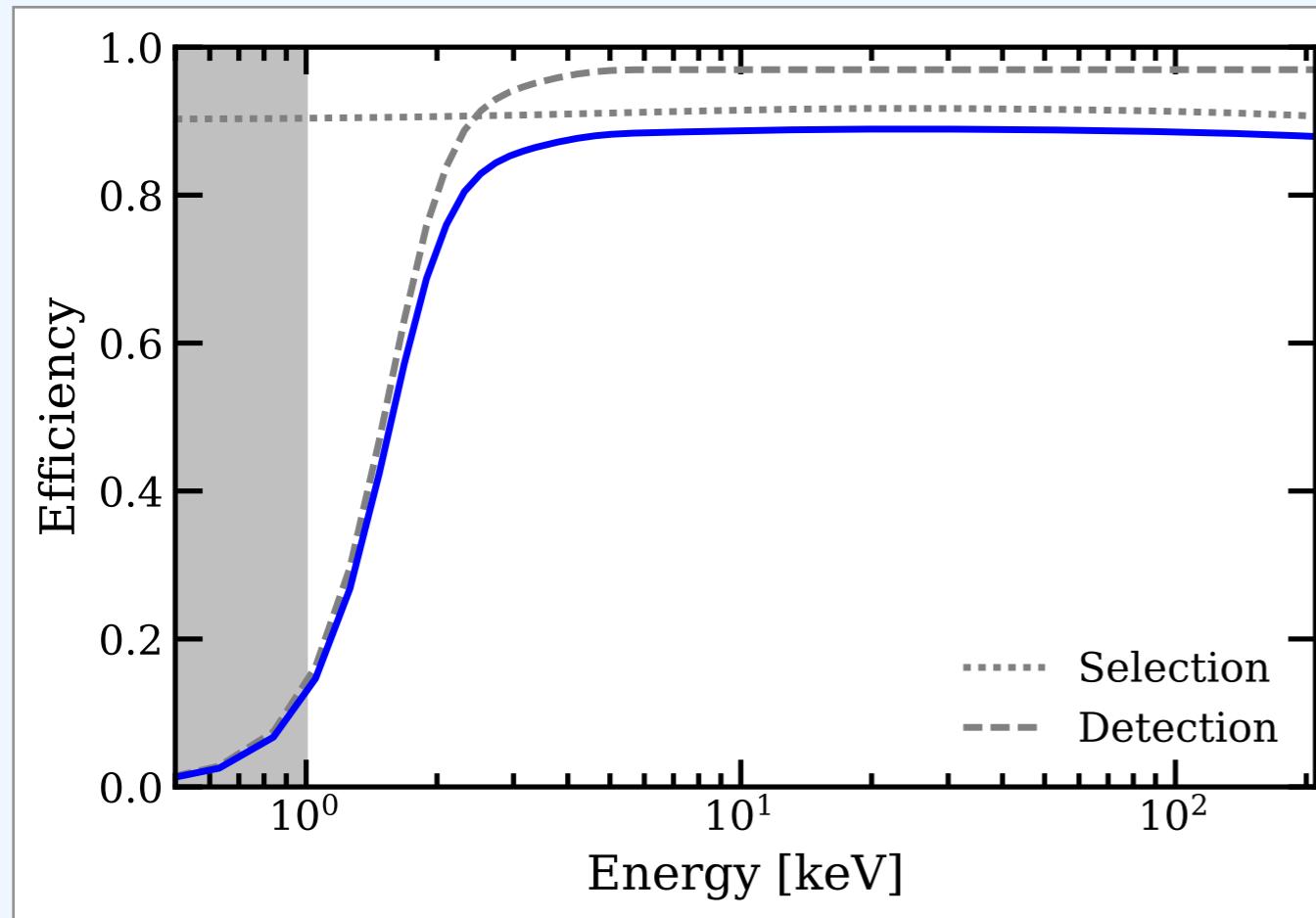


Background Fit

<i>Component</i>	<i>Expected Events</i>	<i>Fitted Events</i>	<i>Constant in time? (shared across partitions)</i>
^{214}Pb	(3450, 8530)	7480 +/- 160	YES
^{85}Kr	890 +/- 50	773 +/- 80	NO
^{136}Xe	2120 +/- 210	2150 +/- 120	YES
^{133}Xe	3900 +/- 410	4009 +/- 85	NO
^{131}Xe	23760 +/- 640	24270 +/- 150	NO
^{83m}Kr	2500 +/- 250	2671 +/- 53	NO
Materials	323 (fixed)	323 (fixed)	YES
Solar neutrino	220.7 +/- 6.6	220.8 +/- 4.7	YES
^{124}Xe	KK	125 +/- 50	YES
	KL	38 +/- 15	YES
	LL	2.8 +/- 1.1	YES
^{125}I	K	79 +/- 33	NO
	L	15.3 +/- 6.5	NO
	M	3.4 +/- 1.5	NO

unconstrained in the fit

Efficiency

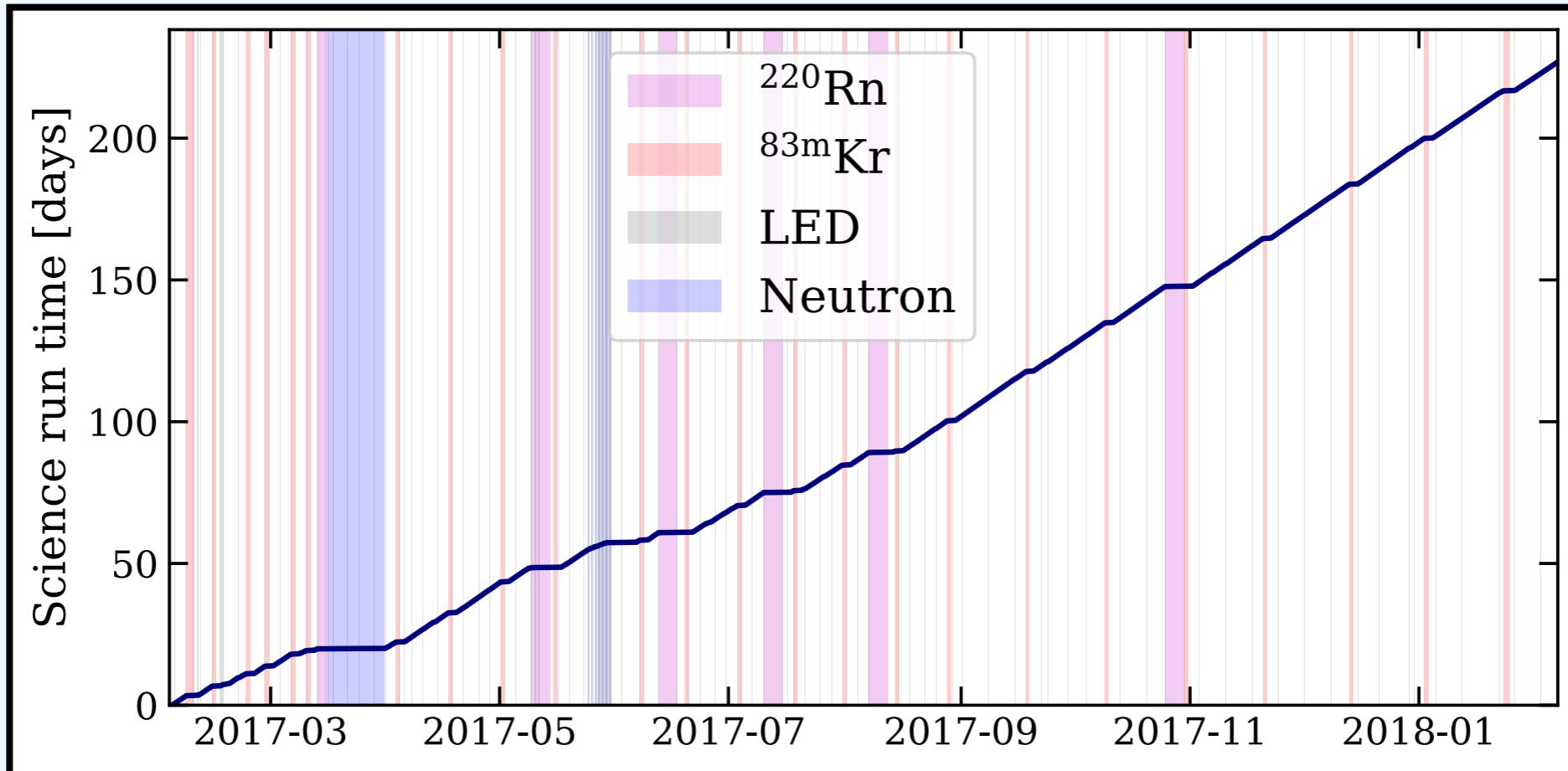


arXiv:1906.04717

3-fold PMT coincidence

Efficiency validated with both waveform simulations and data-driven method using Rn₂₂₀

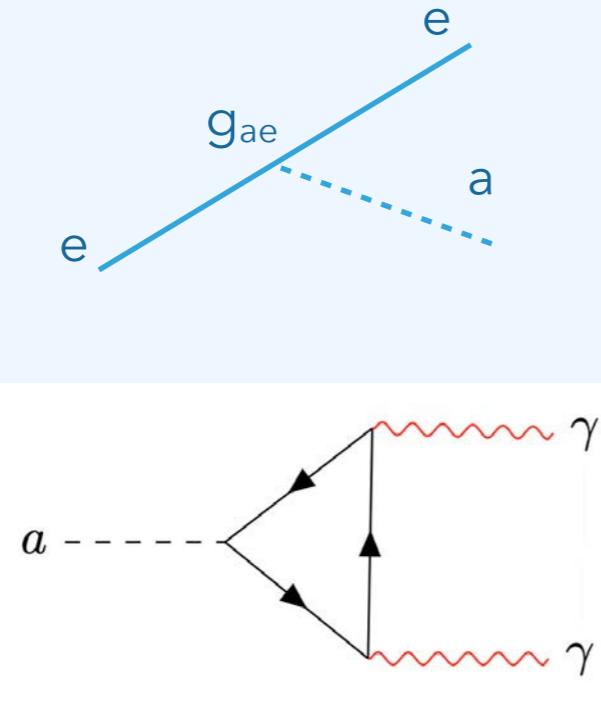
Calibration



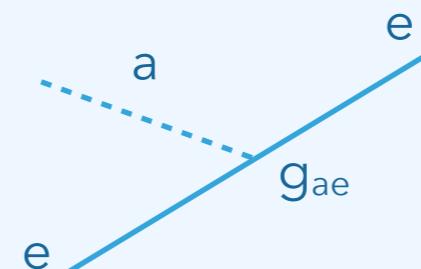
Solar axions



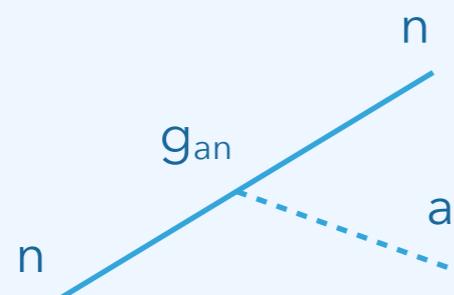
Production
Solar physics



Detection:
Axioelectric effect

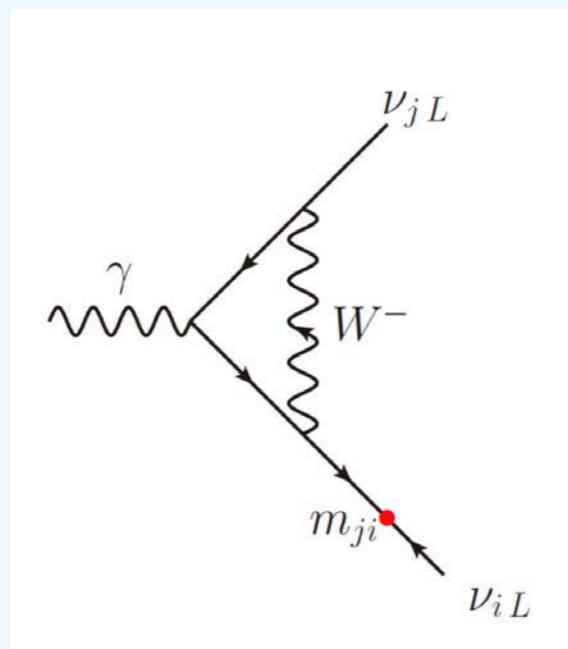


σ for detection, analogous to
photoelectric effect

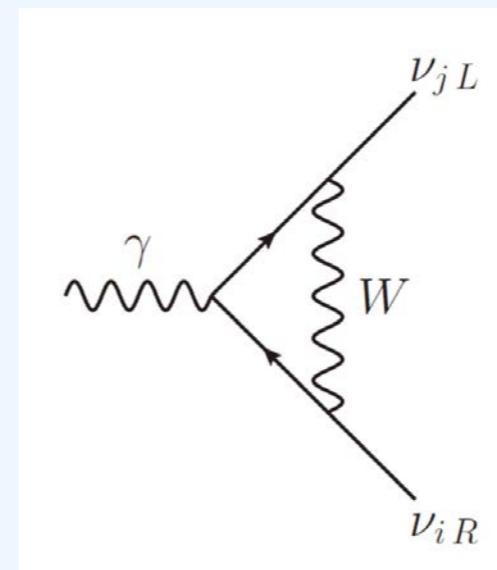


$$\sigma_{ae} = \sigma_{pe} \frac{g_{ae}^2}{\beta} \frac{3E_a^2}{16\pi\alpha m_e^2} \left(1 - \frac{\beta^{2/3}}{3}\right)$$

Neutrino magnetic moment



Majorana neutrinos



Dirac neutrinos

*Jihn E. Kim, 1911.06883, 2019, Bell et al., PLB 642, 2006

The XENON Collaboration



XENON Technical Meeting, May 12-14, 2020

Andrii Terliuk (MPIK/Uni He...)

Alexey Elykov

Ethan Brown

Christopher Hils (JGU-Mai...)

Michele Iacovacci