

# New Physics in the Neutrino Sector

Joachim Kopp

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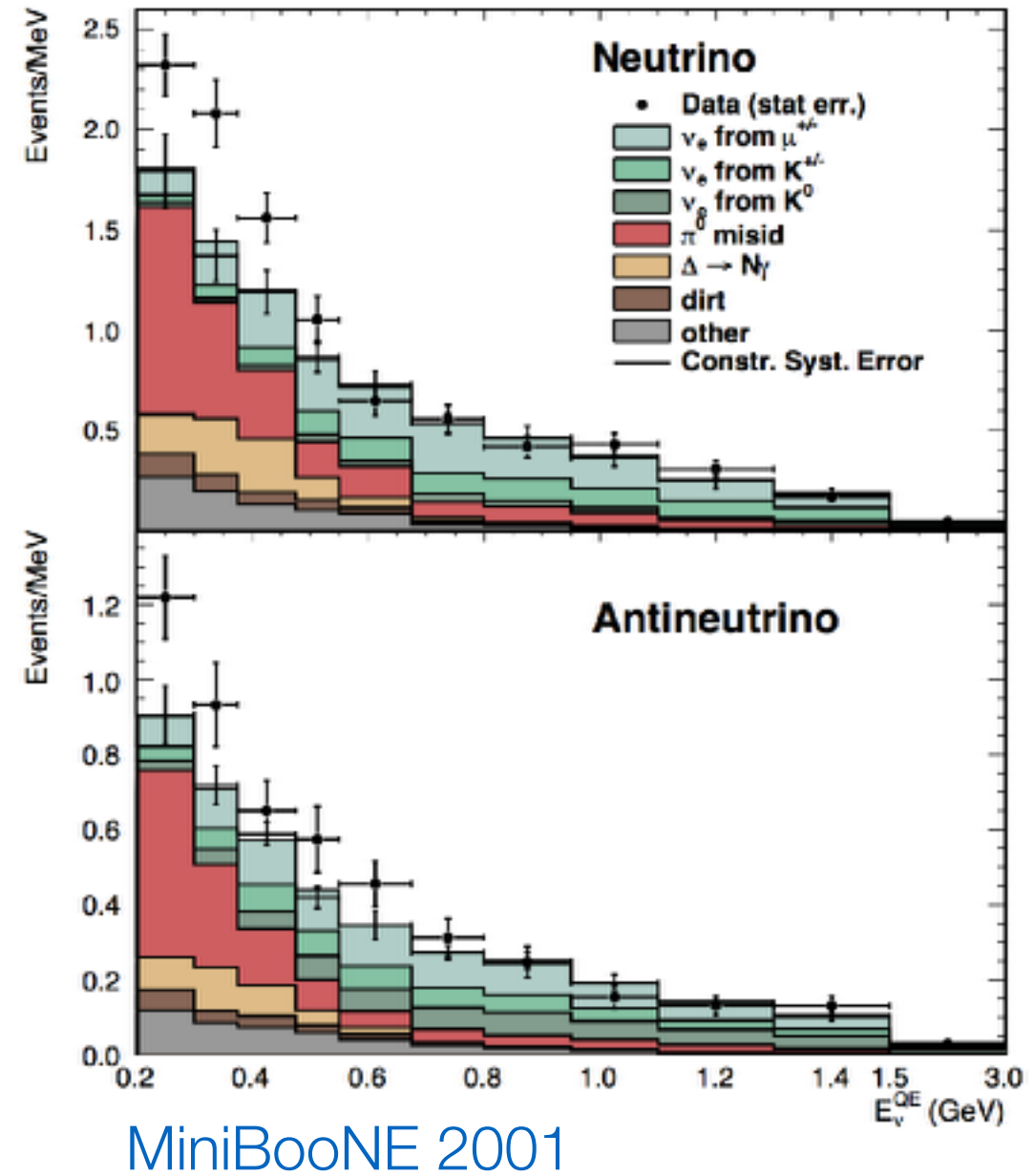
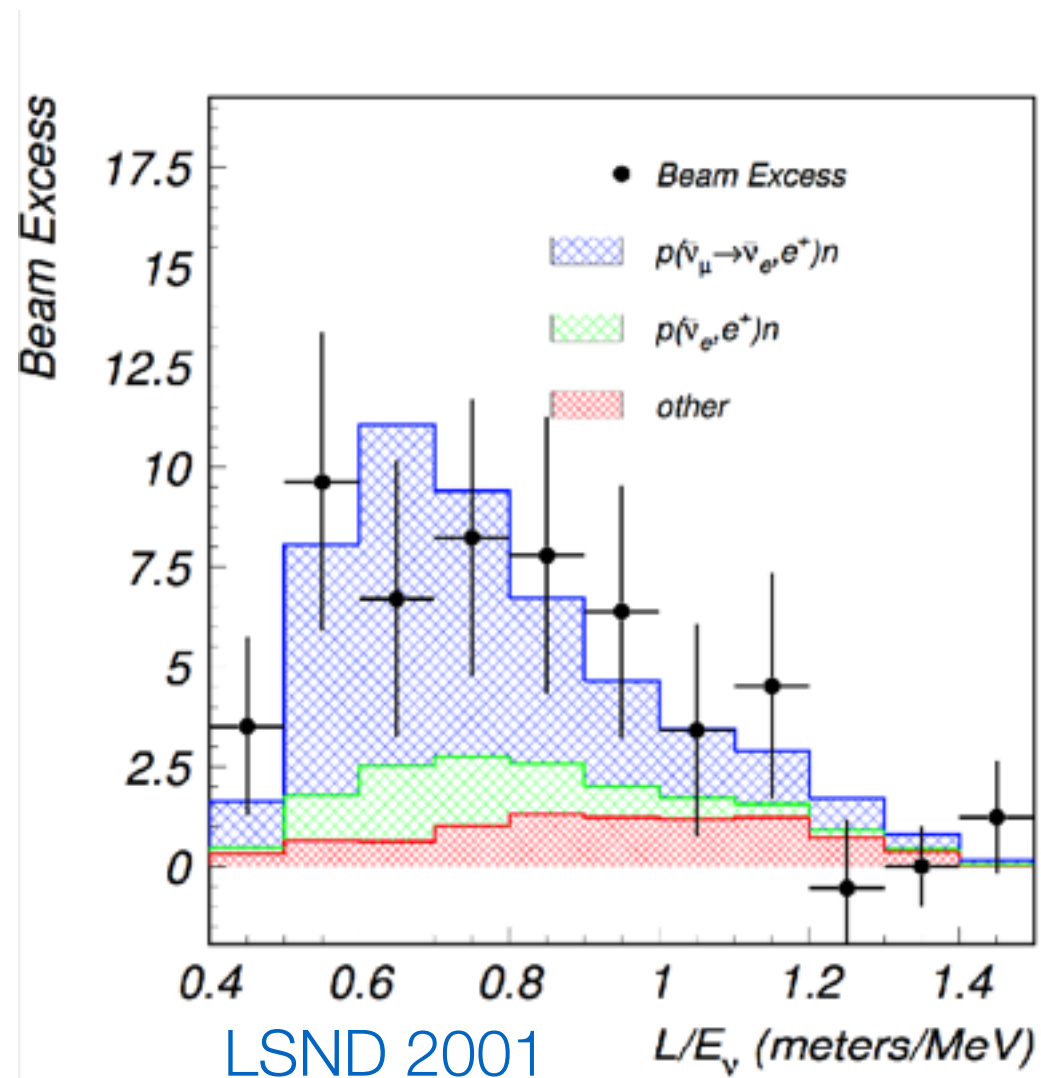


# Light Sterile Neutrinos

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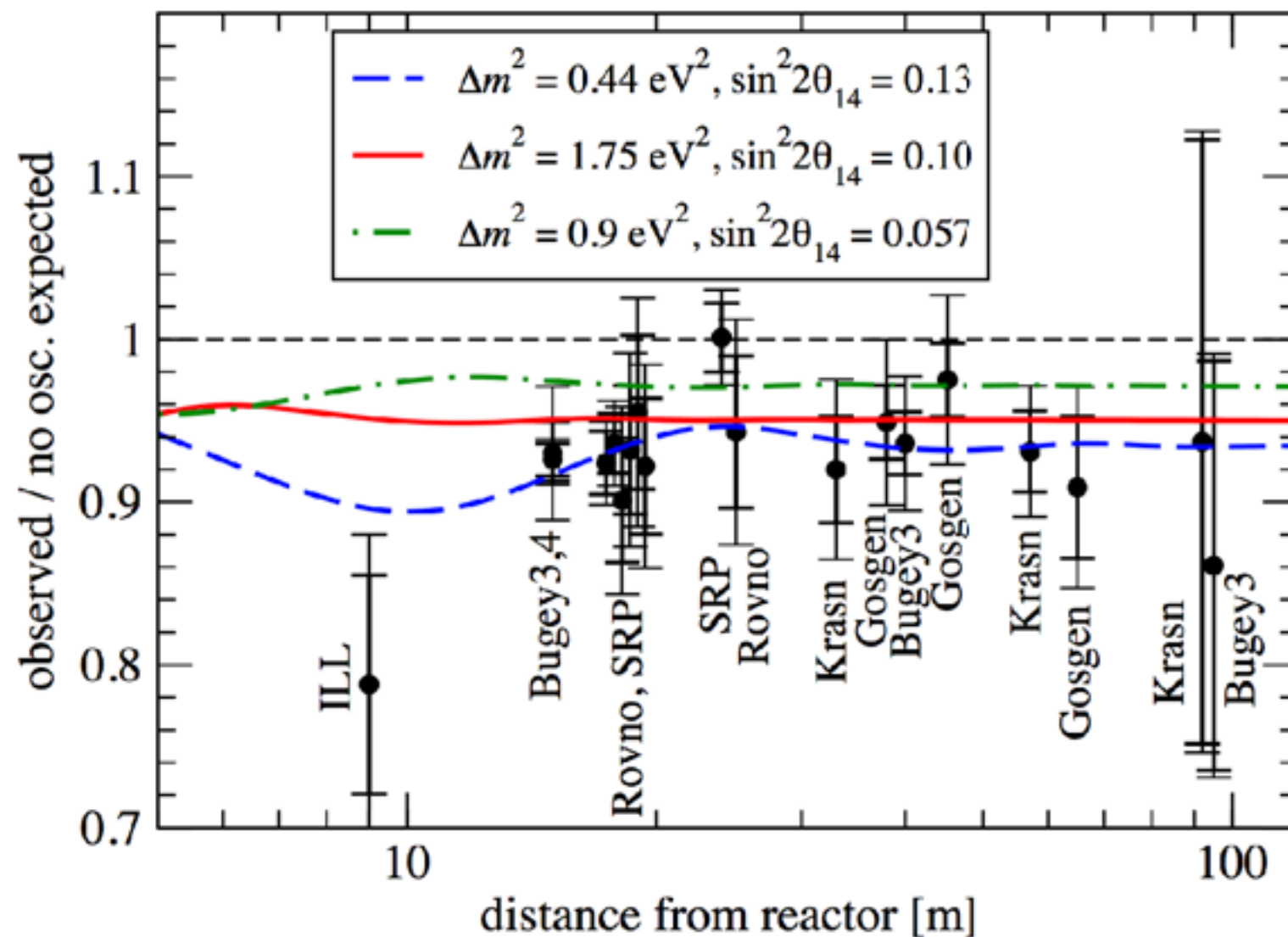


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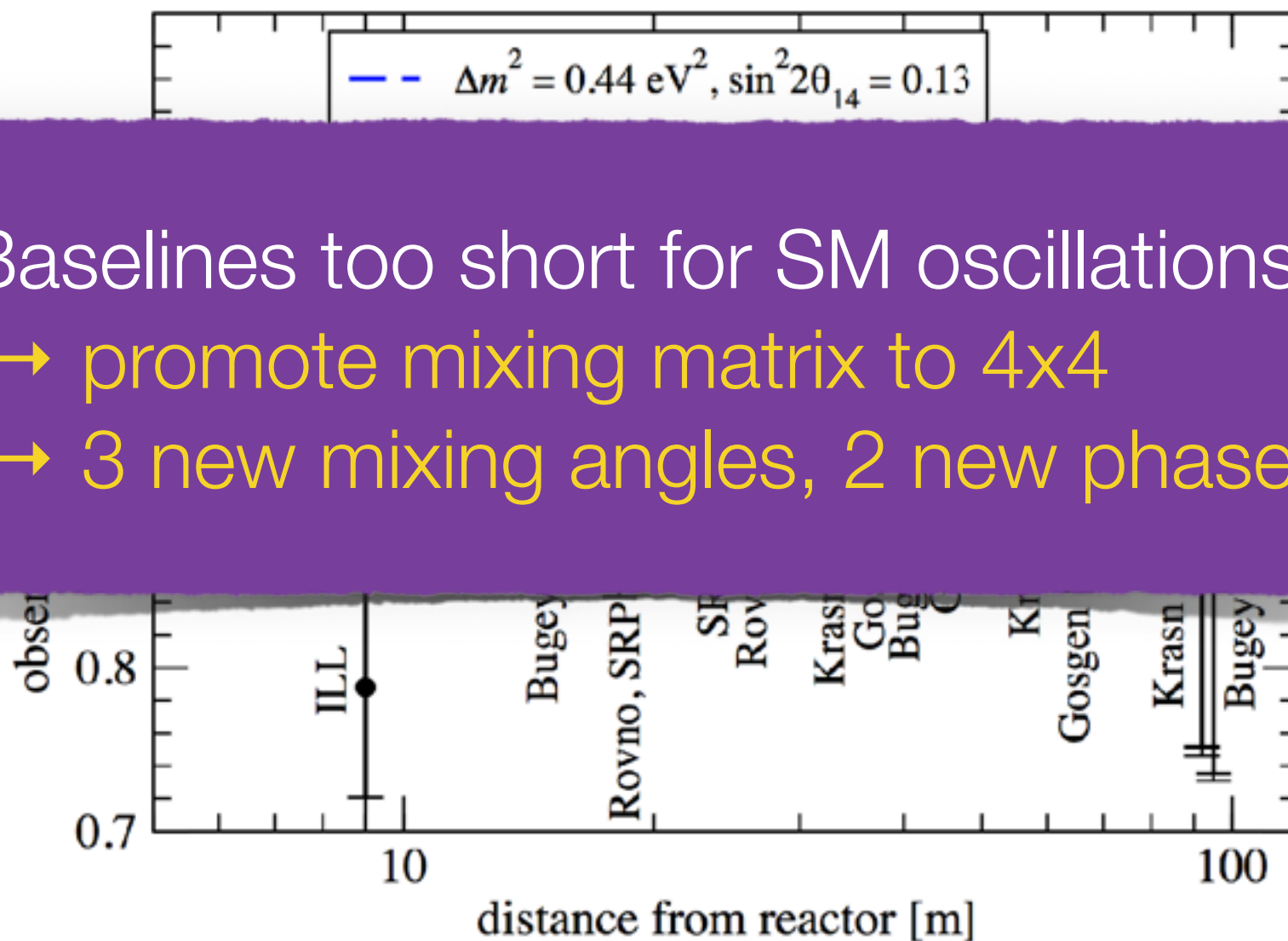


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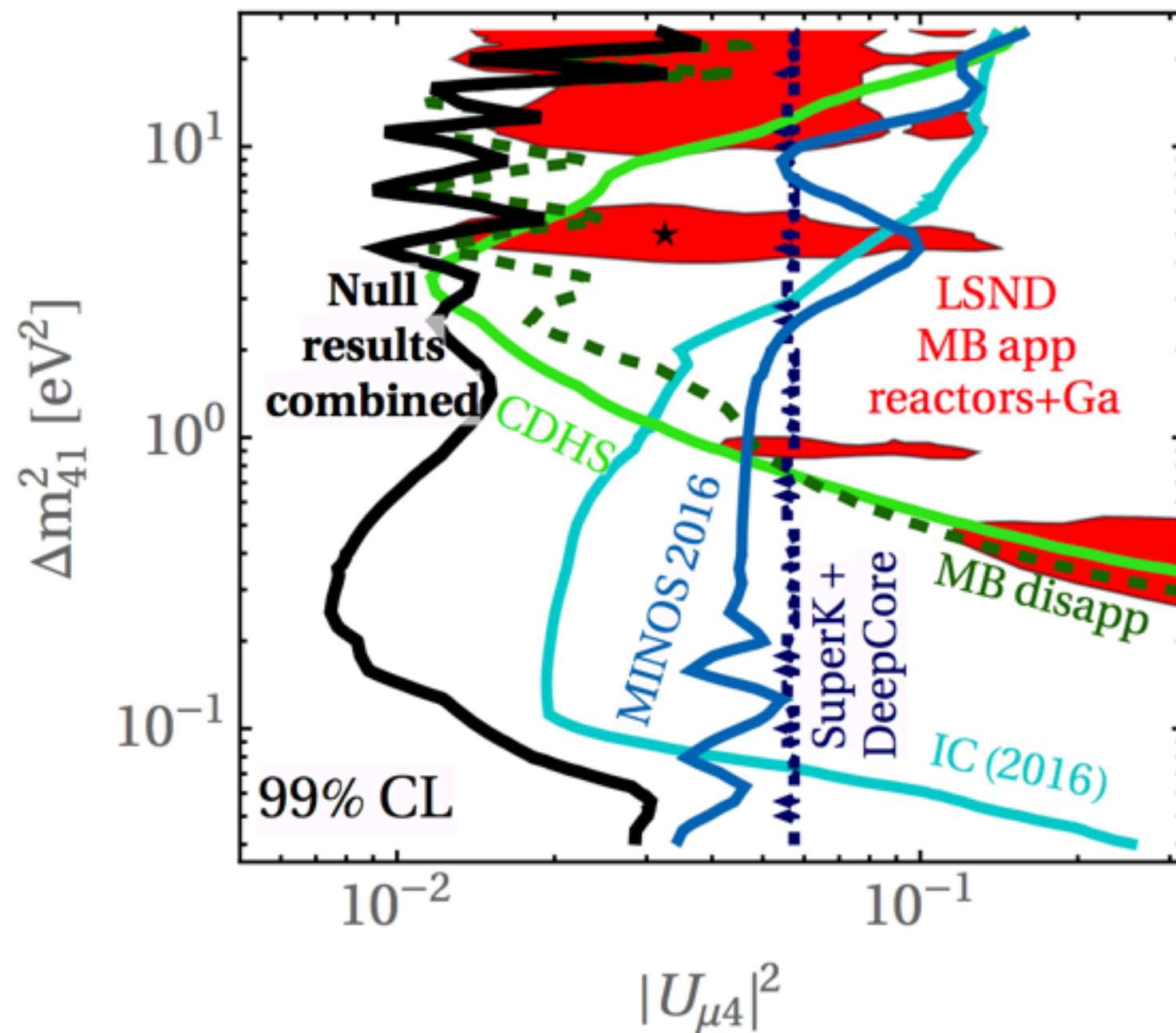
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Baselines too short for SM oscillations!

- promote mixing matrix to 4x4
- 3 new mixing angles, 2 new phases



# Global Fit in 3+1 Model



Dentler Hernandez JK Machado Maltoni Martinez Schwetz, in preparation  
see also works by Collin Argüelles Conrad Shaevitz, [1607.00011](#),  
Gariazzo Giunti Laveder Li, [1703.00860](#)



# Status of Light Sterile Neutrinos



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**severe tension ( $p < 10^{-4}$ )**

- ★ scrutinize anomalies for **unknown systematics**  
(need 4 independent effects!)
- ★ **scrutinize also null results!**

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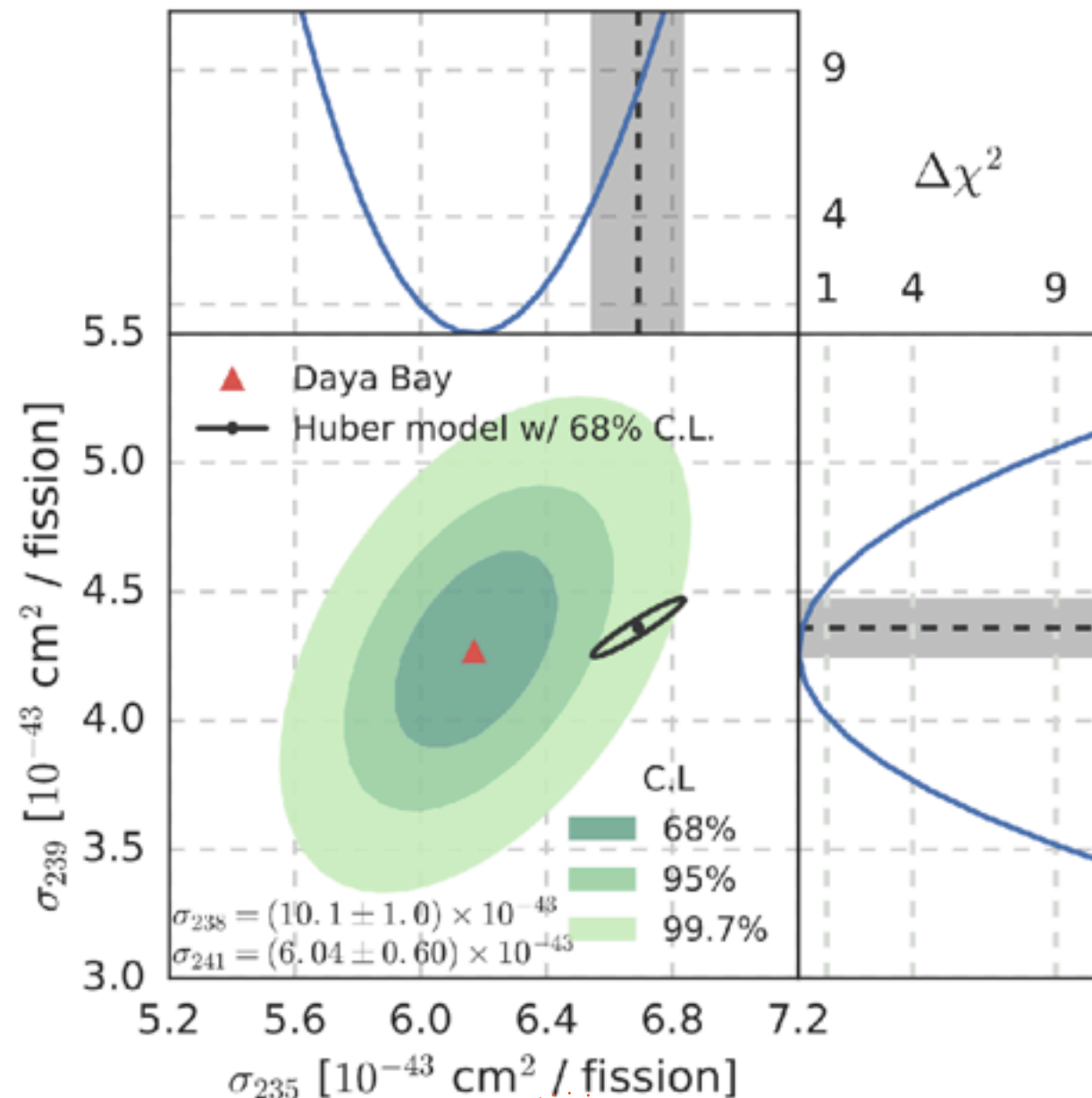
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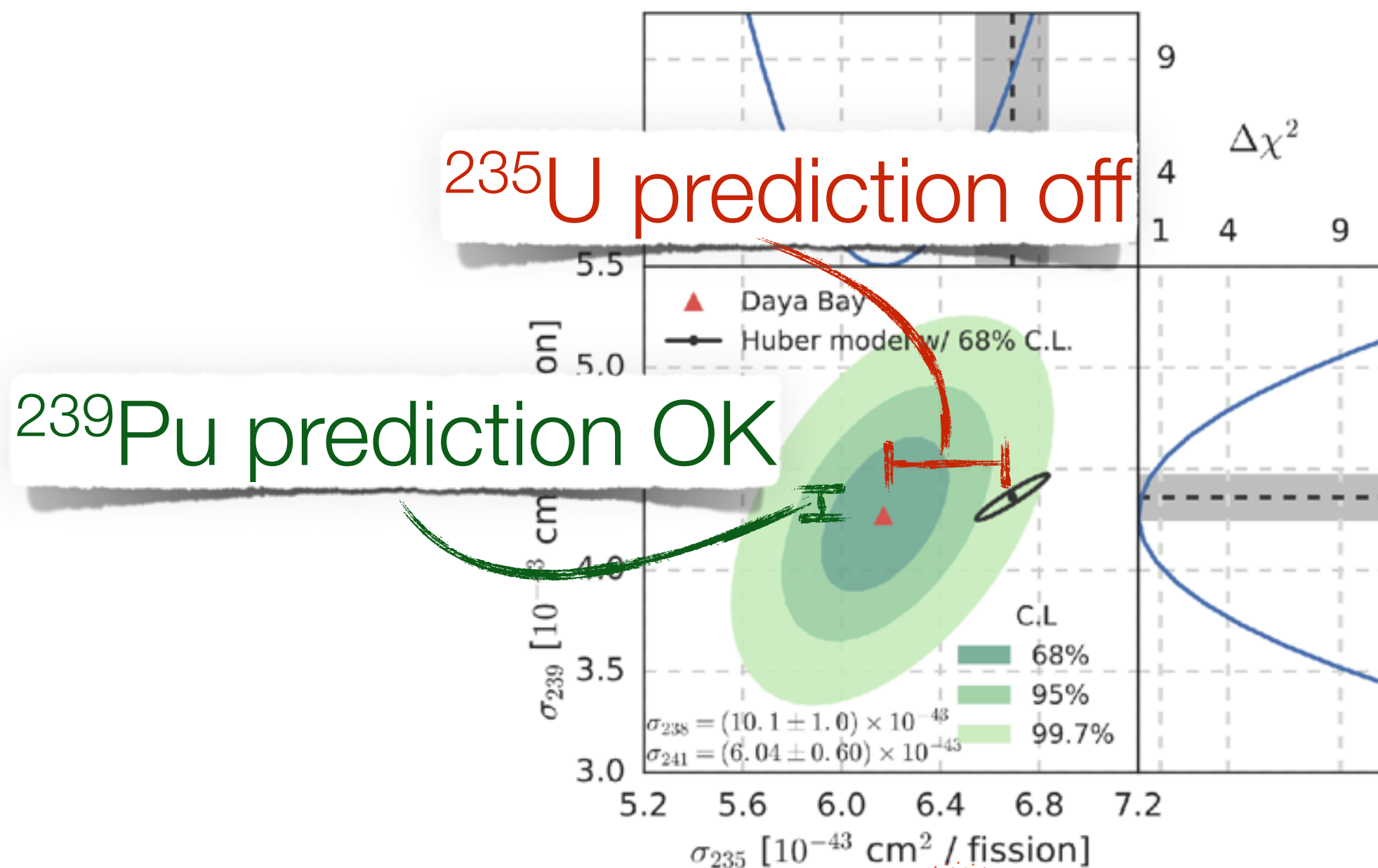
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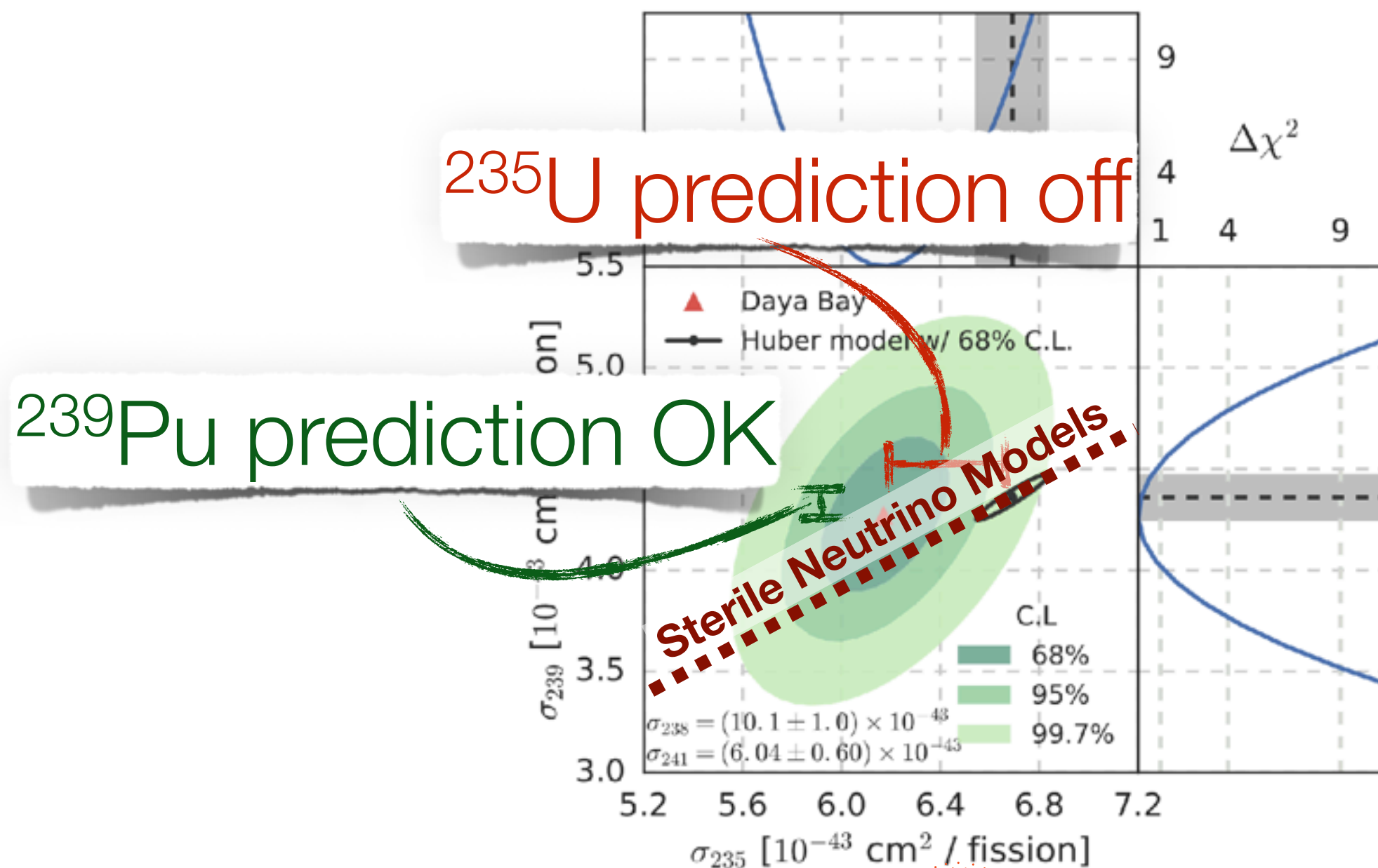
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JK Maltoni Schwetz  
[arXiv:1709.04294](https://arxiv.org/abs/1709.04294)

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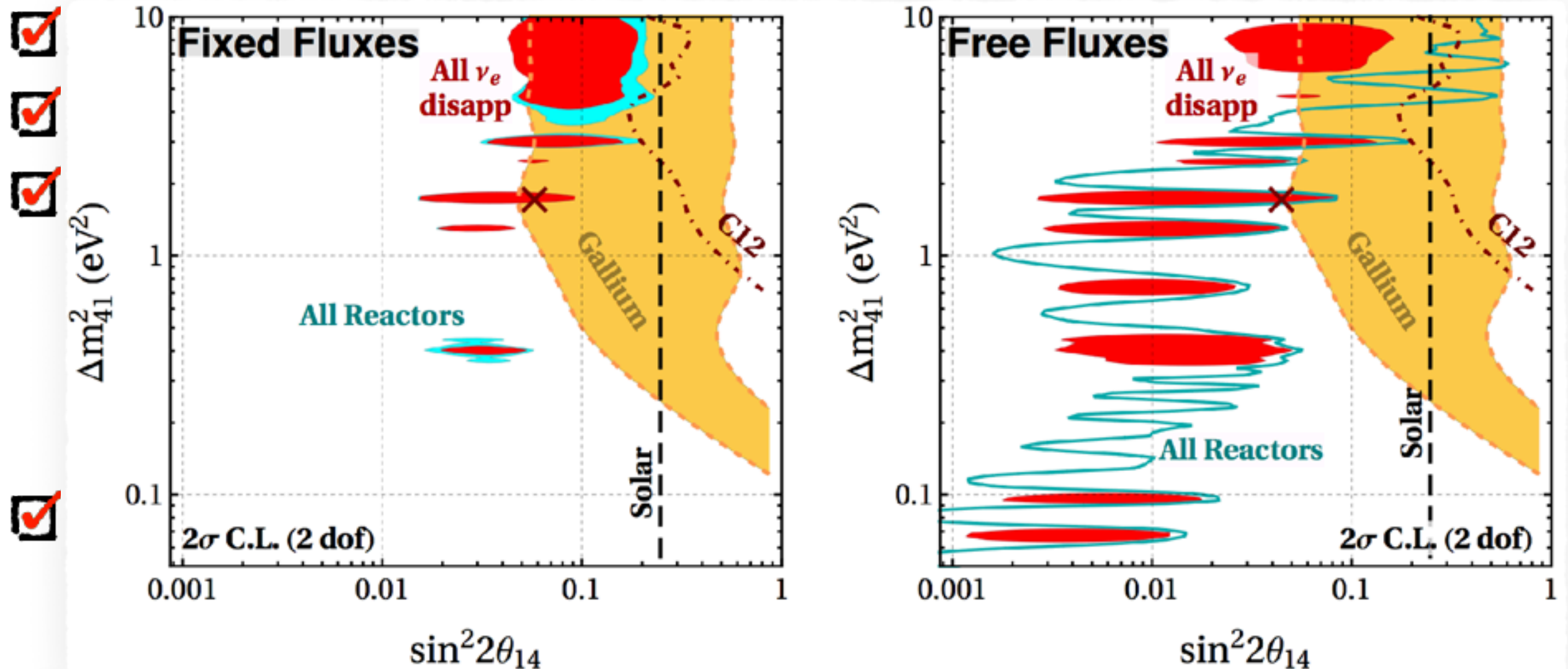
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## ☑ New interactions in the $\nu_s$ sector

- production suppressed by thermal potential  
Hannestad et al. [1310.5926](#); Dasgupta JK, [1310.6337](#)
- minimal scenario now disfavored  
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## ☑ $\nu_s$ properties change in **late phase transition**

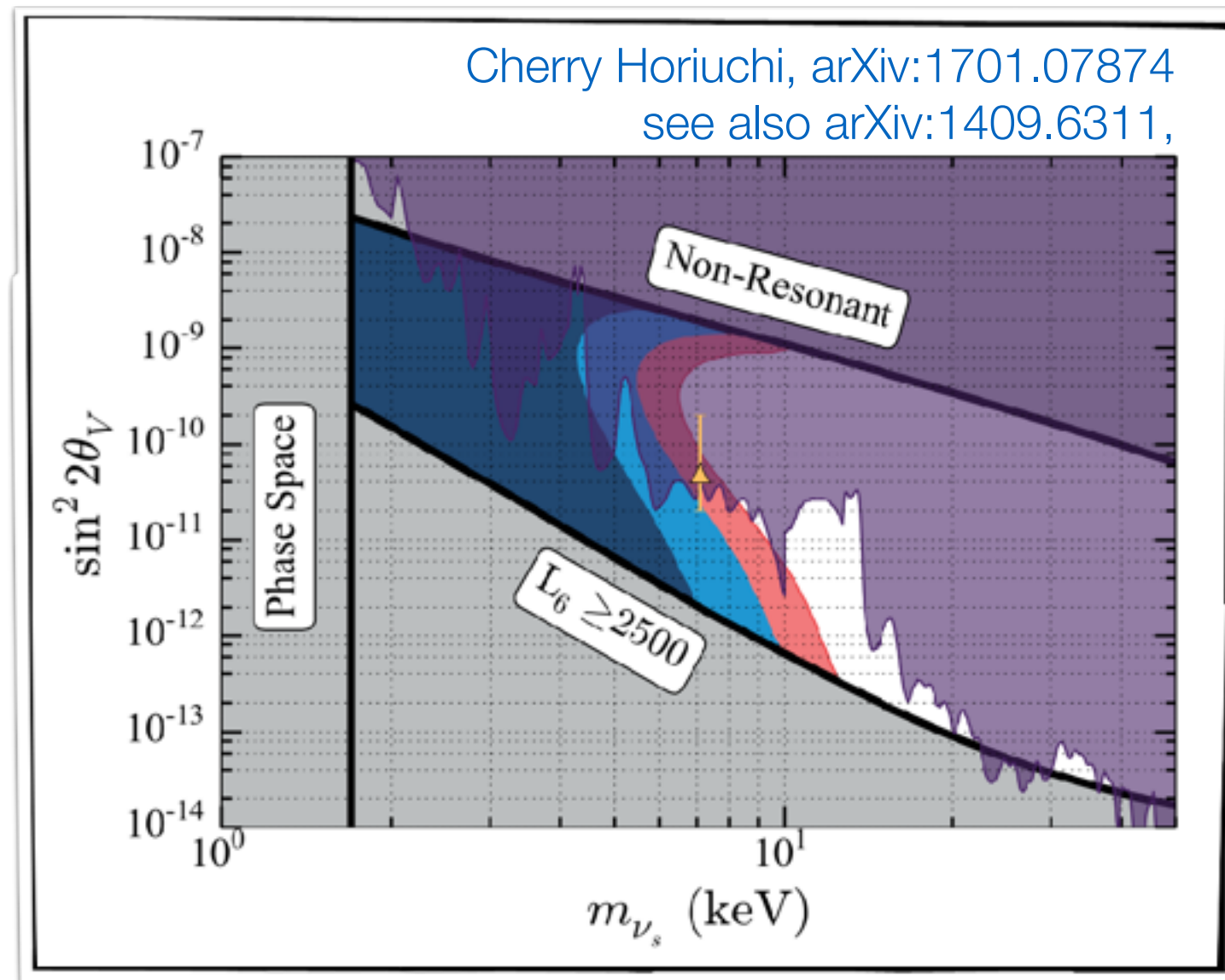
[Bezrukov Chudaykin Gorbunov, 1705.02184](#)

# Neutrinos and Dark Matter

## Recent Developments

# Sterile Neutrinos as Dark Matter

- ☑ keV-scale sterile neutrinos are leading candidate for **Warm Dark Matter**
  - Improved **small scale structure**
  - **x-ray line** signature
- ☑ Production through oscillations challenged by e.g. **Lyman- $\alpha$  data**





# Dark Matter Model Building Flowchart





# Production via Scalar Decays

## ☑ Decay of heavy scalars

$$\mathcal{L} \supset i\overline{N}_\alpha \not{\partial} N_\alpha + \frac{1}{2}(\partial_\mu S)(\partial^\mu S) - \frac{y_\alpha}{2} S \overline{N}_\alpha^c N_\alpha + 2\lambda(H^\dagger H)S^2$$

☑  $S$  freezes in via  $hh \leftrightarrow SS$ , decays via  $S \rightarrow NN$

☑  $N$  produced with relatively cold spectrum

Shaposhnikov Tkachev [hep-ph/0604236](#)

Kusenko [hep-ph/0609081](#)

Merle Niro Schmidt [1306.3996](#)

Merle Totzauer [1502.01011](#)

König Merle Totzauer [1609.01289](#)

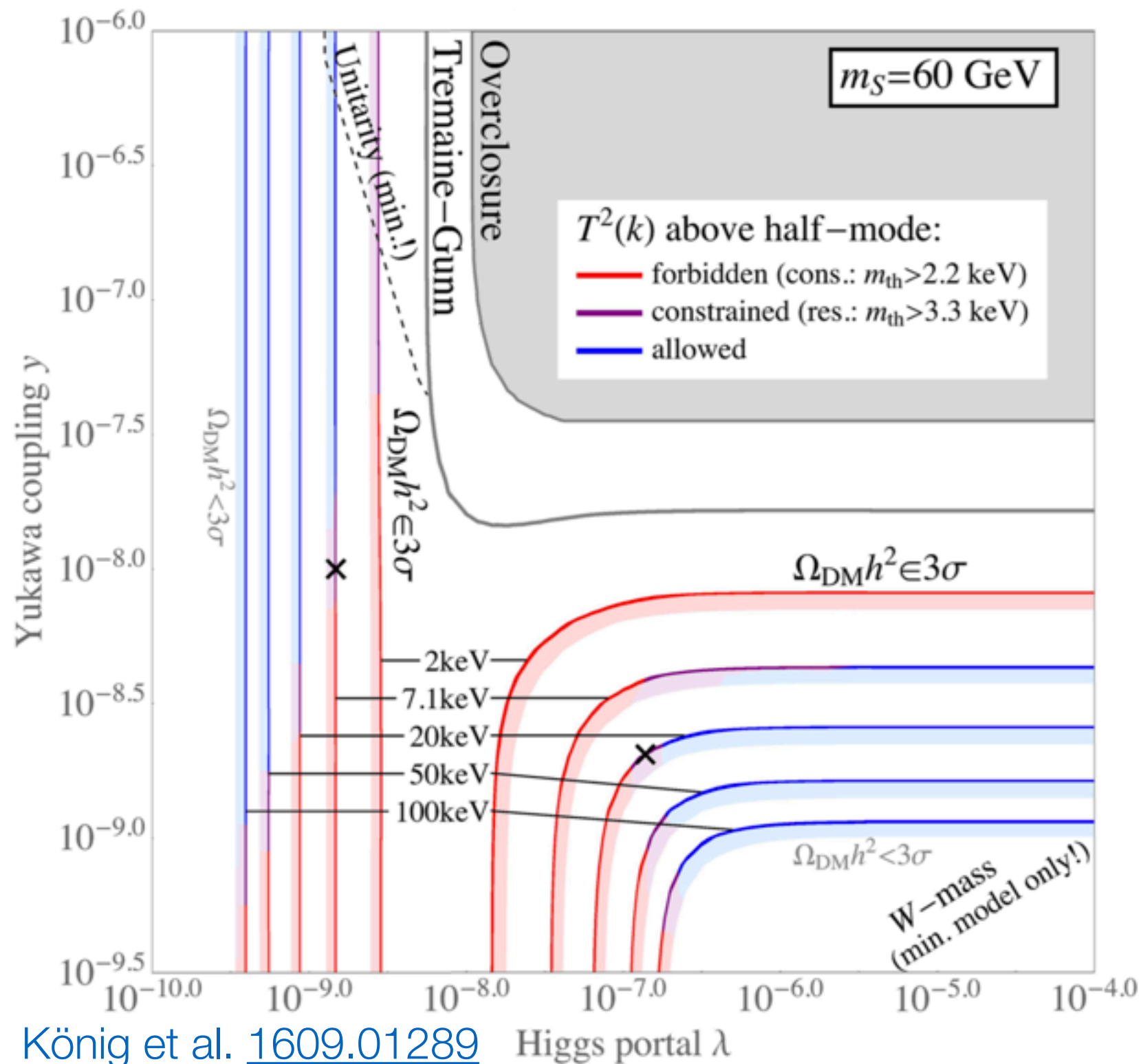
# Production via Scalar Decays

✓ Decay o

$$\mathcal{L} \supset i\overline{N}_c$$

✓  $S$  freeze

✓  $N$  produ



$$H^\dagger H) S^2$$

[hep-ph/0604236](#)  
[hep-ph/0609081](#)  
 dt [1306.3996](#)  
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# Neutrino—DM Interactions

## ☑ Coherent forward scattering of neutrinos on DM

- analogous to SM matter effects (“MSW effect”)
- Observability requires huge DM number density

## ☑ Fuzzy Dark Matter

- scalar or vector,  $m < 10^{-20}$  eV
- Compton wave length  $\sim$  pc
- Interesting for small scale structure

Krnjaic Machado Necib, [1705.06740](#)  
Brdar JK Liu Prass Wang, [1705.09455](#)



## Modified Oscillation Probabilities

### ☒ Coherent

☐ anal

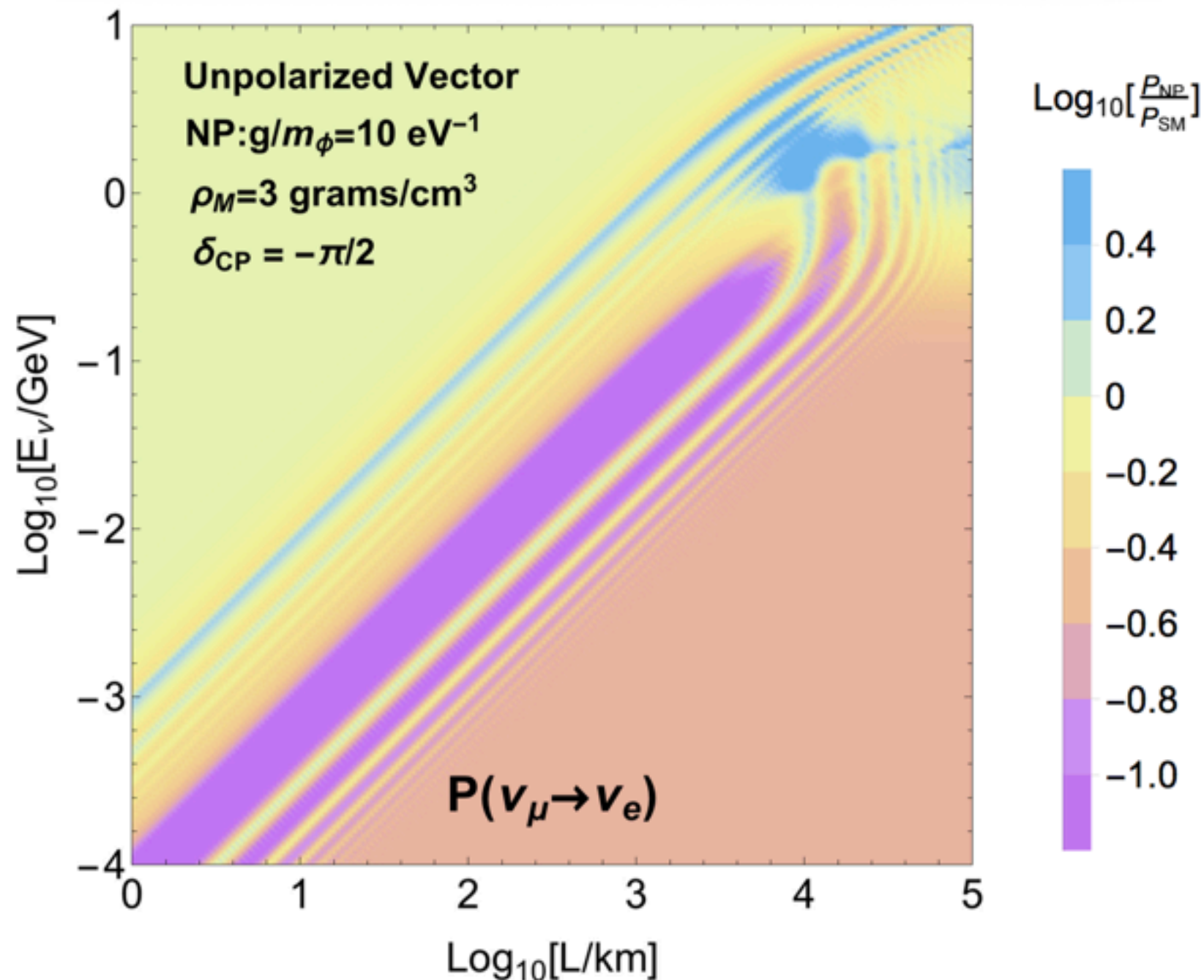
☐ Obs

### ☒ Fuzzy

☐ scal

☐ Con

☐ Inter



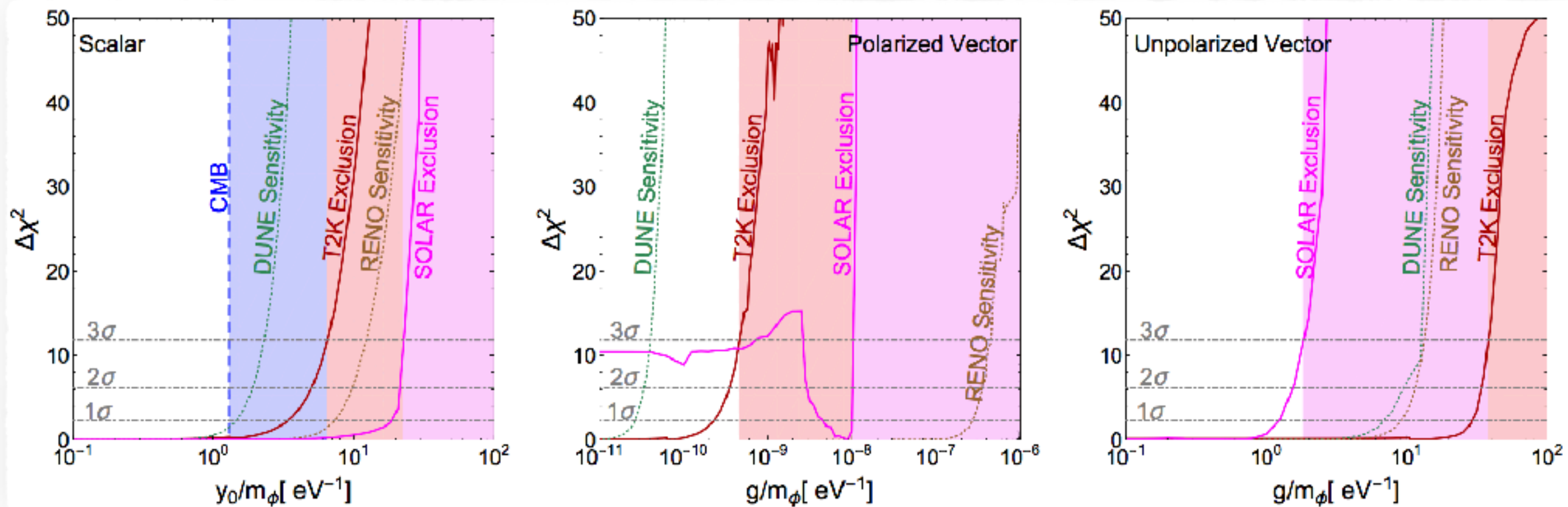
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○ and  $\sigma_{\text{coh}} \approx 10^{-38} \text{ cm}^2 \left( \frac{m_\phi}{1 \text{ GeV}} \right)^4 \left( \frac{g}{10^{-10}} \right)^2$

○ Observed DM relic density  $\Omega_{\text{DM}} h^2 \approx 0.12$

Limits from Long-Baseline Experiments



# Questions?



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